Gauge-Higgs unification in Randall Sundrum Spacetime at Finite Temperature

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Electroweak phase transition and thermal effects

• Our universe : baryon asymmetric (no anti-proton in our daily life)

- Baryogenesis Sakhalov's three conditions
 - **1** B violation process
 - **2** C and CP symmetry is broken
 - 3 out of thermal equilibrium

 Electroweak baryogenesis - the 3rd condition requires the first-order phase transition and the expanding bubbles (inside : broken phase)



Phase transition @high-T for SM and other Models

- Standard Model 2nd order
- 2 SUSY 1st order for some models [Funakubo et.al.]
- **3** Little Higgs Symmetry non-restoration [Espinosa, etl.al.(2004), Aziz, et.al.(2009)]
 - Attempts to get EWPT [Ahriche, et.al. (2009-)]
- In GHU (Hosotani mechanism, flat ExD) 1st order [Ho-Hosotani (1990), Panico, et.al.(2005)]
- 5 UED

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Gauge-Higgs unification (GHU)

• Hierarchy Problem \Leftarrow quadratic divergent δm_h^2

$$m_h^2 = m_{\text{bare}}^2 + (\dots \sim g\Lambda^2), \quad \Lambda : \text{ cutoff}$$

$$m_h = \mathcal{O}(100 \text{GeV}), \quad m_{\text{bare}}, \Lambda \sim M_{GUT} \gg m_h$$
(2)

 \rightarrow fine-tuning between $m_{\rm bare}$ and Λ .

- Gauge-Higgs unification [N.S.Manton (1983), ...]:
 - extra-dimensional component of the gauge field = the Higgs field

$$A_M = (A_\mu, A_y = h) \tag{3}$$

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- gauge symmetry is spontaneously broken by nonzero $\langle A_y \rangle$
- Effective potential and the Higgs-mass is finite, thanks to the gauge symmetry in the higher-dimensional spacetime
 - \rightarrow solve the fine-tuning problem [Inami-Lim-HH (1998)]

Introduction

GHU on Randall Sundrum space-time

- Fermions in S^1/Z_2 extra space:
 - zero-mode wave-function : domain-wall profile due to bulk mass term
 - Yukawa-coupling : overlap of wave-functions of fermions and gauge zero modes:

$$H(x)\bar{\psi}_R(x)\psi_L(x) = \int dy\bar{\psi}_R(x,y)A_y(x,y)\psi_L(x,y)$$
(4)



→ lightest-mode mass depends exponentially on the bulk mass parameter!
 Higgs effective potential (and Higgs mass) are enhanced [Hosotani et.al,2007, HH 2007].

$$m_h \sim \mathcal{O}(100 \text{GeV})$$

(5)

Randall-Sundrum space-time

non-factorizable metric:

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$$ds^2 = e^{-2k|y|} \eta_{\mu\nu} dx^{\mu} dx^{\nu} - dy^2, \quad k : AdS_5 \text{ curvature}$$
(6)

• circle with identification : $y \to -y$ fundamental region : $[0, \pi R]$ fixed points : $y_0 = 0$, $y_1 = \pi R$



Hierarchy

1 UV (hidden brane) scales : Λ, M_5, k, R

2 IR (visible brane) scales : $m_{KK} = \pi k e^{-kR\pi} \frac{1}{1 - e^{-\pi kR}}$

3
$$kR \simeq 12 \rightarrow e^{kR\pi} \simeq M_{\text{Planck}}/M_{\text{Weak}}$$

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Introduction

Finite-Temperature Effects with non-periodic KK tower

• 1-loop effective potential (per field degrees of freedom) at temperature T with Kaluza-Klein mass m_ℓ :

$$V_{\text{eff}}^{1-\text{loop}} = \frac{(-1)^{2\eta}}{2} \frac{1}{\beta} \int \frac{d^3 p}{(2\pi)^3} \sum_{n=-\infty}^{\infty} \sum_{\ell} \ln\left[\left(\frac{2\pi(n+\eta)}{\beta}\right)^2 + \vec{p}^2 + m_{\ell}^2\right],$$

$$\eta = 0(\text{boson}), \frac{1}{2}(\text{fermion}), \quad \beta \equiv 1/T.$$
(7)

• When the extra dimension is compactified on S^1 (radius R),

$$m_{\ell}^2 = \left(\frac{2\pi\ell+ heta}{2\pi R}\right)^2 + M^2, \quad M : \text{bulk mass}$$
 (8)

 \rightarrow one may make use of many tricks (Poisson sum formula, etc...)

 For non-periodic KK modes (e.g. warped compactification) we needs another way of summation.

Introduction

Poisson re-summation only for Matsubara modes gives

$$V_{\text{eff}}^{1-\text{loop}} = V_{\text{eff}}^{T=0} + 2(-1)^{2\eta-1} \sum_{\ell} \sum_{\tilde{n}=1}^{\infty} (-)^{2\eta \tilde{n}} \frac{(\tilde{n}\beta|m_{\ell}|)^2 K_2(\tilde{n}\beta|m_{\ell}|)}{(\sqrt{2\pi}\tilde{n}\beta)^4}$$
(9)

• $x^d K_d(x) \equiv B_d(x)$ is a super-convergent function of x:



 \to By summing up the KK-mode masses from small to large, numerically we can obtain the finite-temperature correction in $V_{\rm eff}.$

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(Review) – SU(2) on $M^3 \times S^1$

[cf. Shiraishi (1987), Ho-Hosotani (1990), Paniko-Serone (2005), Sakamoto-Takenaga (2009)]

Wilson line phase

$$W = \exp(i\theta\sigma_3) = \begin{cases} 1_2 & \theta = 0, \quad SU(2) \text{ unbroken} \\ -1_2 & \theta = \pi, \quad SU(2) \text{ unbroken} \\ \operatorname{diag}(e^{i\theta}, e^{-i\theta}) & \operatorname{otherwise}, \quad SU(2) \to U(1) \end{cases}$$
(10)

• Contributions from fundamental fermion, adjoint fermion, and gauge+ghost fields (blue (cold) \leftrightarrow red (hot))



numerical study

SU(2) on $M^3 \times S^1$ (continued)



GHU in RS

KK modes [Hosotani-Noda etal, 2005]:

$$m_{n} = k\lambda_{n}, \quad G(\alpha, \theta, \lambda_{n}) = 0,$$
(13)

$$G(\alpha, \theta, \lambda_{n}) \equiv \lambda_{n}^{2}z_{1}F_{\alpha-1,\alpha-1}(\lambda_{n}, z_{1})F_{\alpha,\alpha}(\lambda_{n}, z_{1}) - \frac{4}{\pi^{2}}\sin^{2}\frac{\theta}{2},$$

$$\alpha = \begin{cases} \frac{1}{2} \pm (M/k) & \text{fermion} \\ 0, 1 & \text{gauge-ghost-higgs fields} \end{cases}$$
(14)

$$\theta : \text{ Wilson-line phase,}$$

$$F_{\alpha,\beta}(\lambda, z) \equiv Y_{\beta}(\lambda)J_{\alpha}(\lambda z) - J_{\beta}(\lambda)Y_{\alpha}(\lambda z), \quad z_{1} = e^{\pi kR},$$
(15)

$$T = 0) [\text{ Falkowski (2006), HH(2007), Hosotani et.al(2008), Yamashita} \end{cases}$$

• $V_{\text{eff}}(T=0)$ [Falkowski (2006), HH(2007), Hosotani et.al(2008), Yamashita et.al(2008)] $a \equiv e^{-k\pi R}$

$$V^{T=0}(\alpha,\theta) = \frac{k^4 a^4}{16\pi^4} \int_0^\infty dt \, t^3 \operatorname{Re} \ln \left[\frac{G(\alpha,\theta,it)}{G_{\operatorname{asymp}}(\alpha,it)} \right], \quad (16)$$
$$G(\alpha,\theta,it) \stackrel{t \to \infty}{\to} G_{\operatorname{asymp}}(\alpha,it) \quad (17)$$

SU(2) model on RS(preliminary)

• gauge+ghost + 1 fermion (bulk mass M = ck)

- (SU(2) is broken to $U(1)_3$ by orbifold b.c.)
- **gauge** field + fundamental fermion : $U(1)_3$ unbroken
- gauge field + adjoint fermion :



$SO(5) \times U(1)$ GHU model (preliminary)

As application to the particle physics, we study the finite-temperature effect on the model proposed by Hosotani et.al (2008-),

(cf. Hosotani-san's Talk)

- $SO(5) \times U(1) \to SU(2)_L \times \underline{SU(2)_R} \times U(1) \to U(1)_{em}$
- m_h : 70GeV ~ 140GeV
- The model have "H-parity"
 - $P_H = -1$ is assigned for h and +1 for other SM fields
 - All P_H -odd interactions ($hWW, hZZ, h\bar{f}f, hhh$) vanish. → the model can avoid the LEP constraint ($m_h \leq 114$ GeV)
 - *h* is stable and can be the candidate of dark matter (higgs dark-matter).

• Effective potential - we adopt the following approximation:

$$V_{\text{eff}} \simeq \underbrace{V_W + V_Z}_{W,Z\text{boson}} + \underbrace{V_{\text{top}}}_{\text{top guark}},$$
 (18)

$$V_W + V_Z \sim 3V_W = 3 \cdot 3V(1, 2\theta),$$
 (19)

$$V_{\rm top} = -3_{\rm col} \cdot 4V(0.063, 2\theta)$$
 (20)

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other fermions' contribution ($b, c, s, d, u, e, \mu, \tau$, and non-SM heavy particles) are negligible.

- Result for kR = 12 (preliminary)
 - $m_{KK}(\simeq \pi ka) \sim 1.5 \text{TeV}$
 - critical temperature : $\beta_c \sim 1.5/ka \rightarrow T_c \sim 320 \text{GeV}$
 - height of potential barrier : $2 \times 10^{-4} (ka)^4 \sim (60 \text{GeV})^4$
 - $V_c \equiv V_{\text{eff}}(T = T_c) \simeq V_{\text{eff}}(T = 0) \simeq 250 \text{GeV}$
 - \rightarrow unfortunately, not satisfy Shaposhnikov's criteria ($V_c/T_c > 1$)

• kR = 4

- Changes
 - KK scale lowered : $ka \sim 300 \text{GeV} (M_{KK} \sim 1.2 \text{TeV})$
 - top bulk mass for top quark becomes smaller ($c_{\rm top} = [0.294, 0.325]$ for $\pi/2 \le \theta \le \pi)$
- Result for kR = 4 (preliminary)
 - \blacksquare critical temperature : $\beta_c \sim 1.5/ka \rightarrow T_c \sim 200 {\rm GeV}$
 - height of potential barrier : $2 \times 10^{-4} (ka)^4 \sim (50 \text{GeV})^4$
 - $V_c/T_c \ge 1$: satisfy Shaposhnikov's criteria

Image: A mathematical states and a mathem

Summary

Summary

- Numerically studied the Hosotani mechanism on RS at finite-temperature
 - correction from the zero-temperature is obtained by summing up hundreds of Kaluza-Klein masses
 - obtain critical temperature and the height of the potential wall
- Apply to the gauge-Higgs unification ($SO(5) \times U(1)$ model) [preliminary]

	kR = 12	kR = 4
$\exp(\pi kR)$	$2.4 imes 10^{16}$	$2.9 imes 10^5$
m_{KK} , $[ka]$	1.5TeV , [$\sim 500 \text{GeV}$]	1.2TeV, $[\sim 300 \text{GeV}]$
c_{top}	0.437	0.294 - 325
T_c	$330 { m GeV}$	200 GeV
	$(V_c/T_c < 1)$	$(V_c/T_c > 1)$
$V_{\rm barrier}$	$\sim (60 {\rm GeV})^4$	$\sim (50 {\rm GeV})^4$
order of PT	1st	1st

Perspective

- Spharelon process in higher-dimensional space-time
- Flavor mixing, CP violation phase in GHU

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Backup Slides



Figure: Bulk mass parameter for top quark c_{top} versus warp-index kR. Solid (dashed) line is for $\theta_H = \pi/2$ ($\theta_H = \pi$).