

# Propagators in nonstationary spacetimes and a nonequilibrium-thermodynamics character of de Sitter space

Masafumi Fukuma (Dept Phys, Kyoto Univ)

@Osaka University, July 1, 2014

work with **Yuho Sakatani** (Seoul Natinal Univ) and  
**Sotaro Sugishita** (Dept Phys, Kyoto Univ)

## [Refs]

**MF-Sakatani-Sugishita**

**[1] “Propagators in de Sitter space”**

**[PRD88 (2013) 024041, arXiv:1301.7352]**

**[2] “Master equation for the Unruh-DeWitt detector  
and the universal relaxation time in de Sitter space”**

**[PRD89 (2014) 064024, arXiv:1305.0256]**

# **0. Introduction**

**Unification of gravity and the Standard Model is one of the most important problems in contemporary physics.**

**-- Standard Model**

**➡  $\phi(x), \psi(x), A_\mu(x)$ : effective fields at low energy**

**-- What is special for gravity:**

**$g_{\mu\nu}(x)$  is not simply an effective field at low energy.**

**It has a thermodynamic character.**

**[Bekenstein, Hawking, Gibbons-Hawking, ...]**

**Examples:**

**-- Stationary BH has entropy prop to its horizon area.**

**[Bekenstein, Hawking]**

**-- Einstein equation can be regarded as an EoS. [Jacobson]**

**-- Entropy density naturally depends on  $g_{\mu\nu}(x)$ . [MF-Sakatani]**

**Today, I will talk about**

**definition of "vacuum" for QFT in curved spacetime**  
**nonequilibrium-thermodynamics character of de Sitter space**

## **PART 1 : Defining a vacuum in a curved spacetime**

**[1] MF-Sakatani-Sugishita [PRD88 (2013) 024041, 1301.7352]**

- 1. Basics of QFT in curved spacetimes (esp. in de Sitter space)**
- 2. Propagators for the instantaneous vacuum in curved spacetime (general theory)**
- 3. Propagators in de Sitter space**

## **PART 2 : Unruh-DeWitt detector in curved spacetime**

**[2] MF-Sakatani-Sugishita [PRD89 (2014) 064024, 1305.0256]**

- 4. Motivation and the result**
- 5. Master equation for Unruh-DeWitt detector (general theory)**
- 6. Nonequilibrium-thermodynamics character of de Sitter space**

## **PART 1 : Defining a vacuum in a curved spacetime**

**[1] MF-Sakatani-Sugishita [PRD88 (2013) 024041, 1301.7352]**

# 0. Basics of QFT in curved spacetimes

[cf. Birrel-Davies]

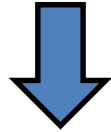
[also 福間-酒谷『重力とエントロピー』(サイエンス社2014)]

## ■ Klein-Gordon inner product

metric:  $ds^2 = g_{\mu\nu}(x) dx^\mu dx^\nu$

action:  $S[\phi(x)] = \int d^d x \left[ -\frac{1}{2} g^{\mu\nu}(x) \partial_\mu \phi(x) \partial_\nu \phi(x) - \frac{m^2}{2} \phi^2(x) \right]$

Sol  $\equiv \{ \text{c-# solns of KG eqn } (\square - m^2) f(x) = 0 \}$



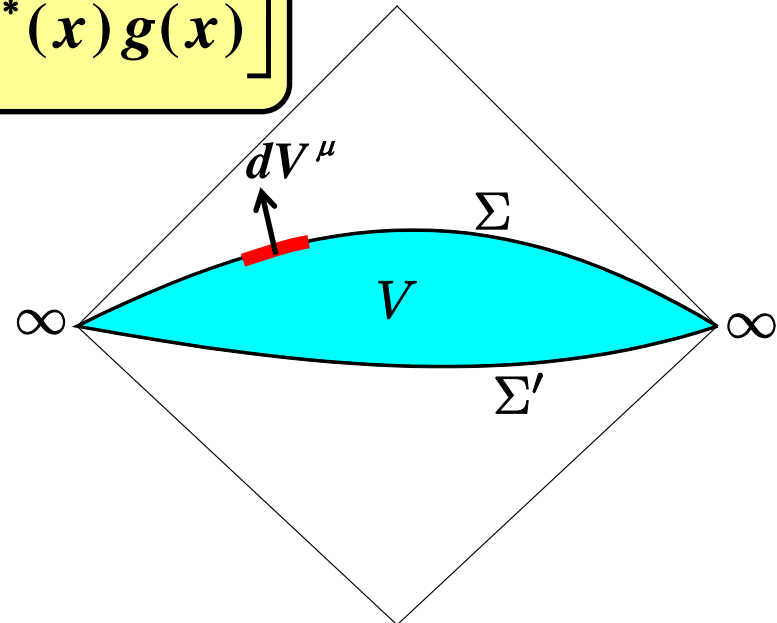
For  $f(x), g(x) \in \text{Sol}$ ,

$$\langle f, g \rangle \equiv i \int_\Sigma dV^\mu \left[ f^*(x) \partial_\mu g(x) - \partial_\mu f^*(x) g(x) \right]$$

“Klein-Gordon inner product”

Properties:

$$\left\{ \begin{array}{l} \text{indep of the choice of } \Sigma \\ \langle g, f \rangle = \langle f, g \rangle^* \\ \langle f^*, g^* \rangle = -\langle g, f \rangle = -\langle f, g \rangle^* \end{array} \right.$$



# ■ Quantization of $\phi(x)$

“wave functions”

## Step 1



Choose a **complete set of positive-energy solns**  $\{\varphi_n(x)\}$ :

maximal set s.t. 
$$\begin{cases} \langle \varphi_n, \varphi_m \rangle = \delta_{nm}, & \langle \varphi_n, \varphi_m^* \rangle = 0, \\ \langle \varphi_n^*, \varphi_m \rangle = 0, & \langle \varphi_n^*, \varphi_m^* \rangle = -\delta_{nm} \end{cases} \quad \text{--- (*)}$$

## Step 2

Expand  $\phi(x)$  as

$$\phi(x) = \sum_n \left[ a_n \varphi_n(x) + a_n^\dagger \varphi_n^*(x) \right] \left( \begin{array}{l} \sum_n \Leftrightarrow \int d^3\mathbf{k} \\ \varphi_n(x) \Leftrightarrow \varphi_{\mathbf{k}}(\mathbf{x}, t) = \frac{e^{i(\mathbf{k}\mathbf{x} - \omega_{\mathbf{k}}t)}}{\sqrt{(2\pi)^3 2\omega_{\mathbf{k}}}} \end{array} \right)$$

## Step 3

Impose the commutation relations

$$\left[ a_n, a_m^\dagger \right] = \delta_{nm}, \quad \left[ a_n, a_m \right] = 0 = \left[ a_n^\dagger, a_m^\dagger \right] \quad \left( \Leftrightarrow \left[ \phi(\mathbf{x}, t), \dot{\phi}(\mathbf{y}, t) \right] = i \delta^3(\mathbf{x} - \mathbf{y}) \right)$$

with (perturbative) **vacuum**  $|0\rangle$ :  $a_n |0\rangle = 0 \quad (\forall n)$



# Quantization of $\phi(x)$

“wave functions”



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## Step 3

**EQUIVALENT**

Impose the commutation relations

$$[a_n, a_m^\dagger] = \delta_{nm}, \quad [a_n, a_m] = 0 = [a_n^\dagger, a_m^\dagger] \quad (\Leftrightarrow [\phi(\mathbf{x}, t), \dot{\phi}(\mathbf{y}, t)] = i \delta^3(\mathbf{x} - \mathbf{y}))$$


with (perturbative) **vacuum**  $|0\rangle$ :  $a_n |0\rangle = 0 \quad (\forall n)$


# ■ Ambiguity in defining a vacuum

## Bogoliubov transformations

$$\varphi'_n(x) = \sum_m \left[ \alpha_{nm} \varphi_m(x) + \beta_{nm} \varphi_m^*(x) \right]$$

$$\text{s.t.} \begin{cases} \alpha \cdot \alpha^\dagger - \beta \cdot \beta^T = 1, & \alpha \cdot \beta^T - \beta \cdot \alpha^T = 0, \\ \alpha^\dagger \cdot \alpha - \beta^T \cdot \beta = 1, & \alpha^\dagger \cdot \beta - \beta^T \cdot \alpha^* = 0 \end{cases}$$

 Relations (\*) hold also for  $\{\varphi'_n(x)\}$ .

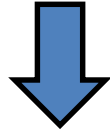
$\exists$  ambiguity of  $\{\varphi_n(x)\}$   ambiguity of  $|0\rangle$

 ambiguity of Feynman propagator:

$$\begin{aligned} G(x, y) &= \langle 0 | \mathbf{T} \phi(x) \phi(y) | 0 \rangle \\ &= \theta(x^0 - y^0) \sum_n \varphi_n(x) \varphi_n^*(y) + \theta(y^0 - x^0) \sum_n \varphi_n(y) \varphi_n^*(x) \end{aligned}$$

## ■ What is a natural vacuum?

case 1:  $\exists$  global timelike Killing vector  $\xi = \xi^\mu(x)\partial_\mu$

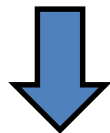


Using  $t$  s.t.  $\xi = \partial_t$ ,

we choose  $\{\varphi_n(x) = \varphi_n(\mathbf{x}, t)\}$  s.t. they take the form

$$\varphi_n(x) = \varphi_n(\mathbf{x}, t) = \varphi_n(\mathbf{x}) e^{-i\omega_n t} \quad (\omega_n > 0)$$

case 2:  $\exists$  vector  $\xi = \xi^\mu(x)\partial_\mu$  that is asympt. timelike Killing



$$\begin{array}{l} \xi \rightarrow \partial_{t_{\text{in}}} \text{ (far past)} \quad \longrightarrow \quad |0; \text{in}\rangle \\ \xi \rightarrow \partial_{t_{\text{out}}} \text{ (far future)} \quad \longrightarrow \quad |0; \text{out}\rangle \end{array} \quad \left( \begin{array}{l} \text{Generally,} \\ |0; \text{in}\rangle \not\propto |0; \text{out}\rangle \end{array} \right)$$

**de Sitter space does not belong to case 1 or 2.**

# **1. Basics of QFT in de Sitter space**

**[1] MF-Sakatani-Sugishita [PRD88 (2013) 024041, 1301.7352]**

## ■ $d$ -dimensional de Sitter space

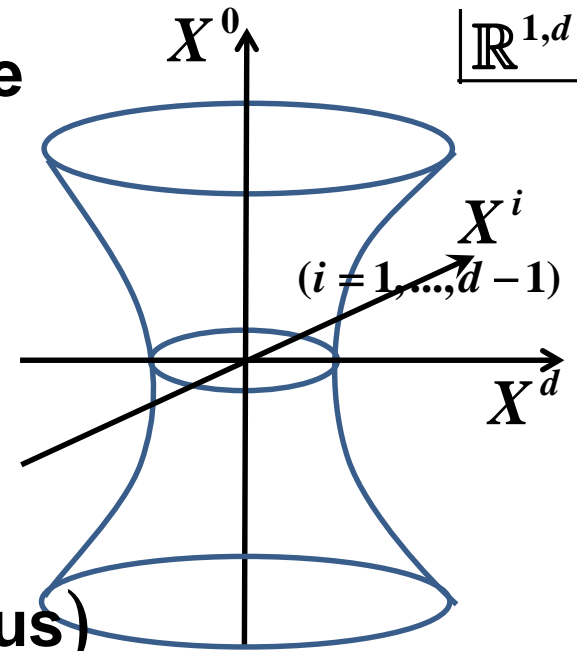
hyperboloid in  $(d + 1)$ -dim Minkowski space

$$-(X^0)^2 + (X^1)^2 + \dots + (X^{d-1})^2 + (X^d)^2 = \ell^2$$

(topology:  $\mathbb{R} \times S^{d-1}$ )

This is a solution to Einstein's equation  
with positive cosmological constant

$$\Lambda = \frac{(d-2)(d-1)}{2\ell^2} > 0 \quad (\ell : \text{curvature radius})$$



## ■ QFT in de Sitter space

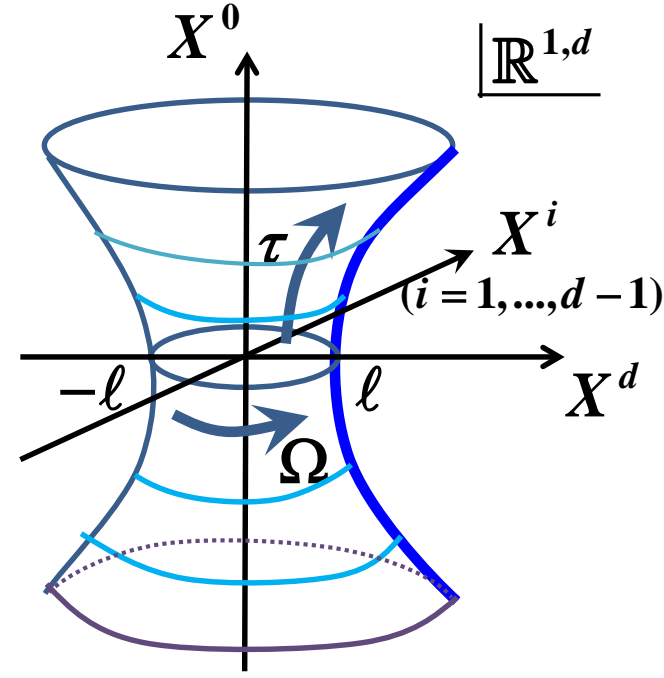
- Motivation
- one of the simplest but nontrivial spacetimes
  - application to  $\left\{ \begin{array}{l} \text{inflation} \\ \text{cosmological const problem} \end{array} \right.$
  - helpful in understanding the thermal character of curved spacetime

# various coordinate patches

## (1) global patch

$$\begin{cases} X^0 = \ell \sinh \frac{\tau}{\ell} \\ X^I = \ell \cosh \frac{\tau}{\ell} \Omega^I \quad (I = 1, \dots, d) \\ \left( \Omega \in S^{d-1}; \quad \sum_{i=1}^d (\Omega^I)^2 = 1 \right) \end{cases}$$

$$ds^2 = -d\tau^2 + \ell^2 \cosh^2 \frac{\tau}{\ell} d\Omega_{d-1}^2$$

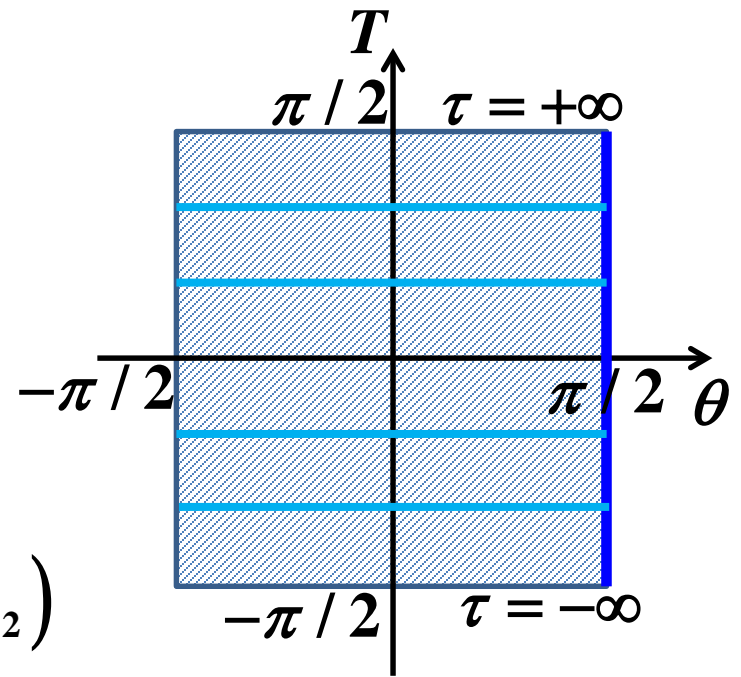


## Penrose

$$\begin{aligned} \sin T &= \tanh \frac{\tau}{\ell}, \\ \Omega^I &= (\Omega^i, \theta) \\ (-\pi/2 < \tau, \theta < \pi/2) \end{aligned}$$

$$\begin{pmatrix} X^0 = \ell \sinh \frac{\tau}{\ell} \\ X^d = \ell \sin \theta \\ X^i = \ell \cos \theta \Omega^i \\ (i = 1, \dots, d-1) \end{pmatrix}$$

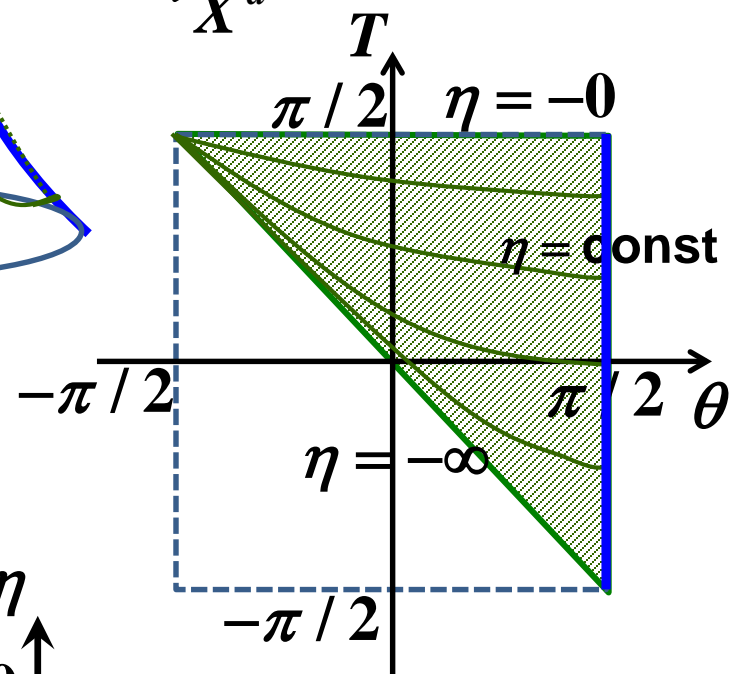
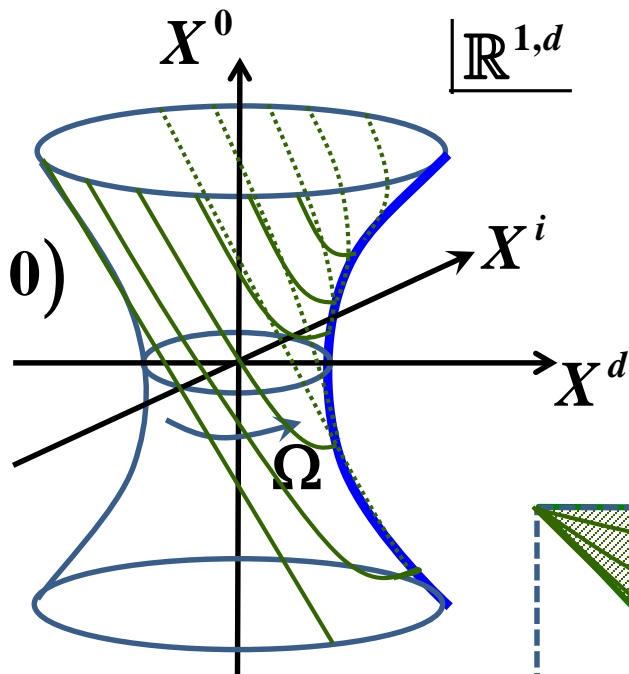
$$\rightarrow ds^2 = \frac{\ell^2}{\cos^2 T} (-dT^2 + d\theta^2 + \cos^2 \theta d\Omega_{d-2}^2)$$



## (2) Poincaré patch

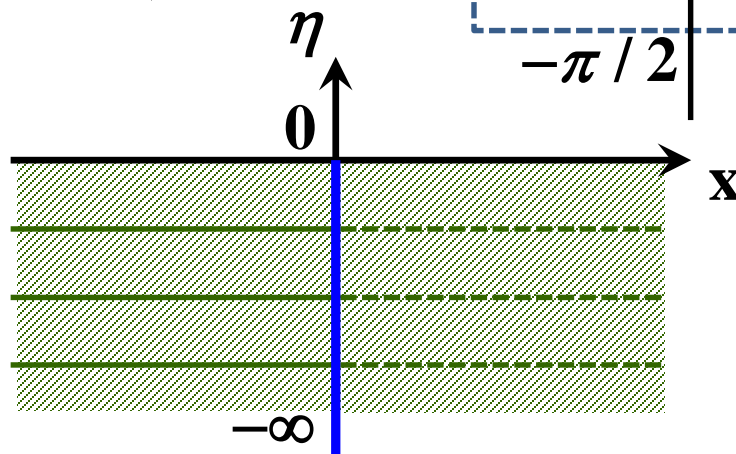
$$\begin{cases} X^0 = \frac{\ell^2 - \eta^2 + \mathbf{x}^2}{-2\eta} \\ X^d = \frac{\ell^2 + \eta^2 - \mathbf{x}^2}{-2\eta} \\ X^i = \frac{\ell x^i}{-\eta} \quad (i = 1, \dots, d) \end{cases}$$

$(\eta < 0)$



$$\begin{aligned} ds^2 &= \ell^2 \frac{-d\eta^2 + d\mathbf{x}^2}{\eta^2} \\ &= -dt^2 + e^{2t/\ell} d\mathbf{x}^2 \end{aligned}$$

$$(-\eta = \ell e^{-t/\ell})$$



**NB:**

$$X^0 + X^d = \frac{\ell^2}{-\eta} > 0$$

## ■ de Sitter invariant

$$\eta_{MN} X^M X^N = -(X^0)^2 + (X^1)^2 + \dots + (X^d)^2 = \ell^2$$

correspondence

$$x \in \mathbf{dS} \Leftrightarrow X(x) \in \mathbb{R}^{1,d}$$



invariant under  $O(1,d)$ :

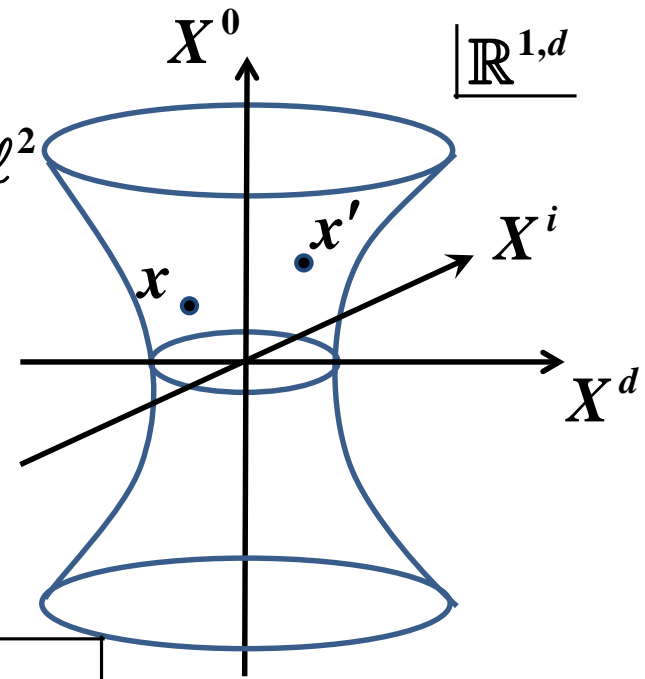
$$Z(x, x') \equiv \frac{1}{\ell^2} \eta_{MN} X^M(x) X^N(x') \equiv \frac{1}{\ell^2} X \cdot X'$$

[global]  $\left( ds^2 = -d\tau^2 + \ell^2 \cosh^2 \frac{\tau}{\ell} d\Omega_{d-1}^2 \right)$

$$Z(x, x') = -\sinh \frac{\tau}{\ell} \sinh \frac{\tau'}{\ell} + \cosh \frac{\tau}{\ell} \cosh \frac{\tau'}{\ell} \Omega \cdot \Omega'$$

[Poincaré]  $\left( ds^2 = \ell^2 \frac{-d\eta^2 + d\mathbf{x}^2}{\eta^2} \right)$

$$Z(x, x') = \frac{\eta^2 + \eta'^2 - |\mathbf{x} - \mathbf{x}'|^2}{2\eta\eta'}$$





## ■ Issue of defining a vacuum in de Sitter space

no global timelike Killing vector



time-dependent Hamiltonian



no stationary states

If adiabaticity holds asymptotically  
(i.e. no level crossing occurs at the far past/future),  
one can define an asymptotic vacuum.

This cannot be expected for de Sitter space  
(except for  $\eta = -\infty$  in Poincare).

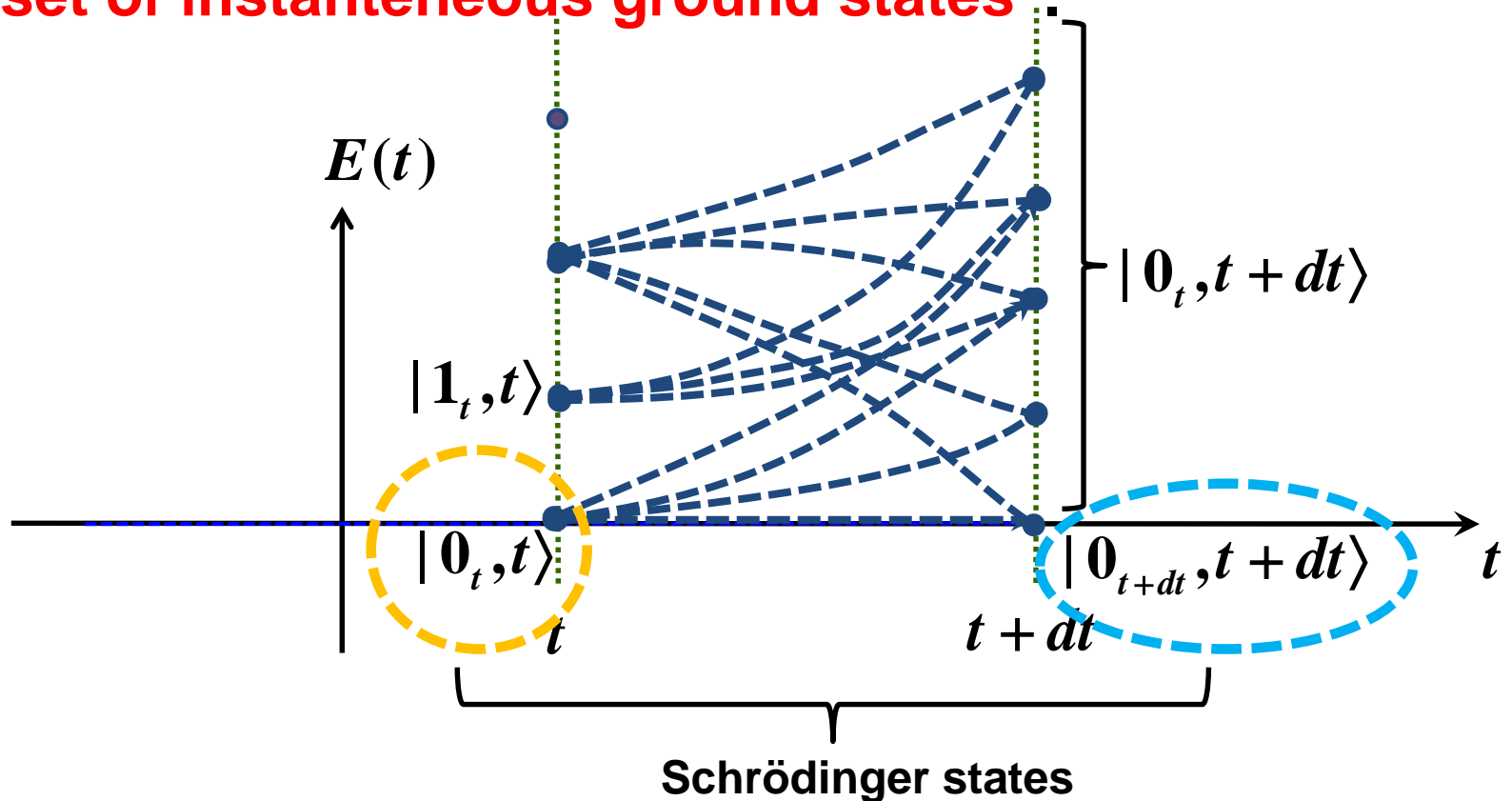


There exists an ambiguity in defining propagators.

■ Our strategy [MF-Sakatani-Sugishita [1] 1301.7352]

- Instead of defining a global vacuum, we consider the ground state at each instant, and define the vacuum to be

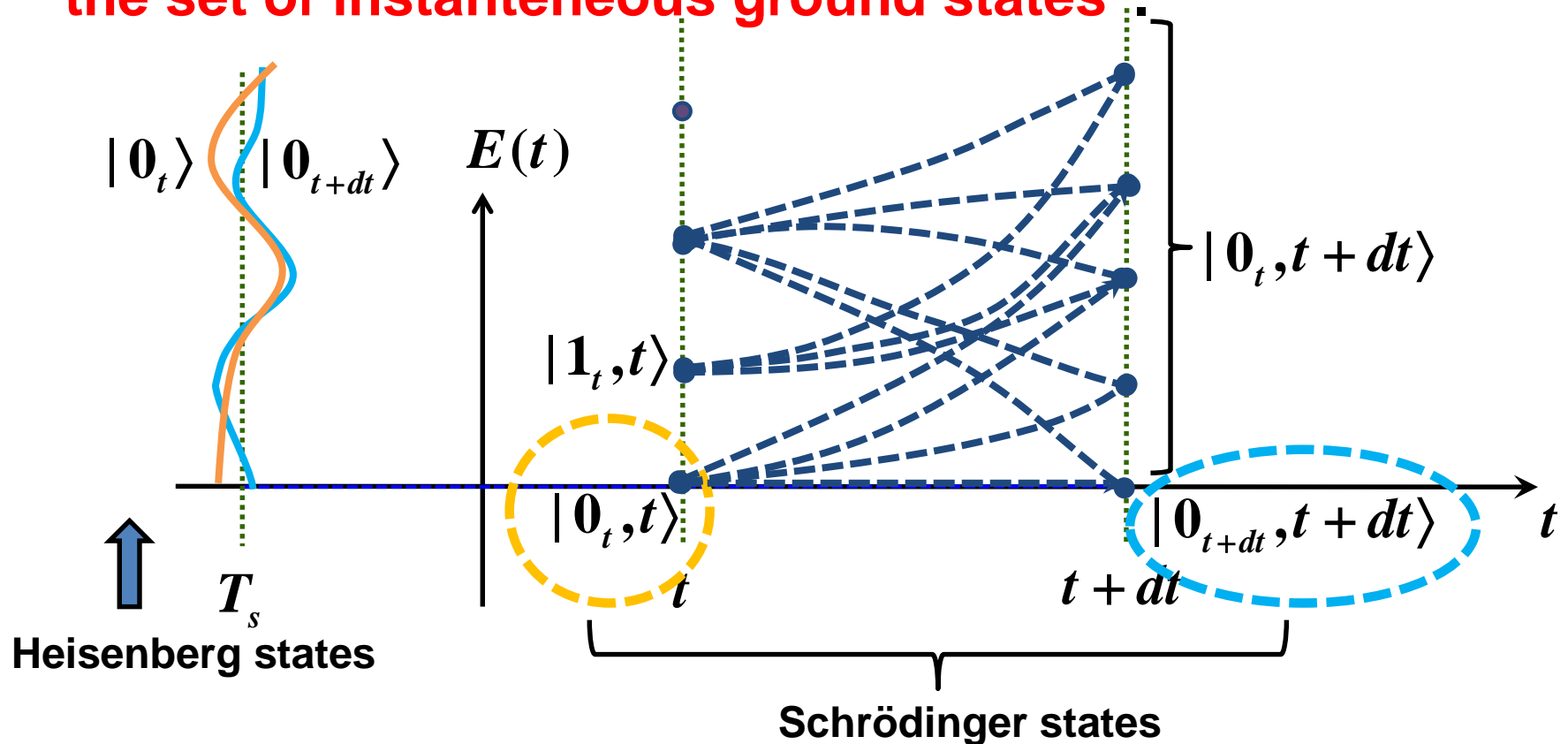
"the set of instantaneous ground states"



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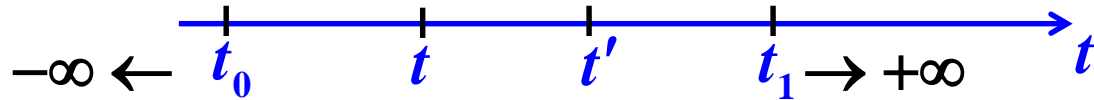
"the set of instantaneous ground states"



- Introduce 2-pt functions:

$$G_{00}(x, x'; t_0, t_0) \equiv \frac{\langle \mathbf{0}_{t_0} | \mathbf{T} \phi(x) \phi(x') | \mathbf{0}_{t_0} \rangle}{\langle \mathbf{0}_{t_0} | \mathbf{0}_{t_0} \rangle}$$

$$G_{10}(x, x'; t_1, t_0) \equiv \frac{\langle \mathbf{0}_{t_1} | \mathbf{T} \phi(x) \phi(x') | \mathbf{0}_{t_0} \rangle}{\langle \mathbf{0}_{t_1} | \mathbf{0}_{t_0} \rangle}$$



- Take the limit  $t_0 \rightarrow -\infty, t_1 \rightarrow +\infty$ :

$$G^{\text{in/in}}(x, x') \equiv \lim_{t_0 \rightarrow -\infty} G_{00}(x, x'; t_0, t_0) \equiv \lim_{t_0 \rightarrow -\infty} \frac{\langle \mathbf{0}_{t_0} | \mathbf{T} \phi(x) \phi(x') | \mathbf{0}_{t_0} \rangle}{\langle \mathbf{0}_{t_0} | \mathbf{0}_{t_0} \rangle}$$

$$G^{\text{out/in}}(x, x') \equiv \lim_{\substack{t_0 \rightarrow -\infty \\ t_1 \rightarrow +\infty}} G_{10}(x, x'; t_1, t_0) \equiv \lim_{\substack{t_0 \rightarrow -\infty \\ t_1 \rightarrow +\infty}} \frac{\langle \mathbf{0}_{t_1} | \mathbf{T} \phi(x) \phi(x') | \mathbf{0}_{t_0} \rangle}{\langle \mathbf{0}_{t_1} | \mathbf{0}_{t_0} \rangle}$$

- Result: A careful calculation shows that the finite limit exists for all patches.  
(except for  $G^{\text{in/in}}$  in the global patch for  $m < (d - 1) / 2\ell$ )

## **2. Propagators for the instantaneous vacuum in curved spacetime (general theory)**

**[1] MF-Sakatani-Sugishita [PRD88 (2013) 024041, 1301.7352]**

## ■ Mode expansion

Assume that the metric takes the form

$$ds^2 = -N^2(t) dt^2 + A^2(t) h_{ij}(\mathbf{x}) dx^i dx^j$$

Introduce a complete set  $\{Y_n(\mathbf{x}) \in \mathbb{R}\}$

for the spatial Laplacian  $\Delta_{d-1} \equiv \frac{1}{\sqrt{h}} \partial_i (\sqrt{h} h^{ij} \partial_j)$  s.t.

$$\Delta_{d-1} Y_n(\mathbf{x}) = -\lambda_n Y_n(\mathbf{x}), \quad \int d^{d-1}\mathbf{x} \sqrt{h} Y_n(\mathbf{x}) Y_m(\mathbf{x}) = \delta_{n,m}$$

Expand

$$\phi(\mathbf{x}) = \phi(t, \mathbf{x}) = \sum_n \phi_n(t) Y_n(\mathbf{x}) \quad (\phi_n(t) \in \mathbb{R})$$



$$S[\phi] = \sum_n \int dt \left[ \frac{\rho(t)}{2} \dot{\phi}_n^2(t) - \frac{\rho(t) \omega_n^2(t)}{2} \phi_n^2(t) \right]$$

with  $\rho(t) = N^{-1}(t) A^{d-1}(t)$  and  $\omega_n^2(t) = N^2(t) [\lambda_n A^{-2}(t) + m^2]$

**“a system of decoupled harmonic oscillators”**

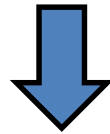
# Hamiltonian (in the Schrödinger picture)

$$H_s(t) = \sum_n H_{n,s}(t) = \sum_n \left[ \frac{\pi_{n,s}^2}{2\rho(t)} + \frac{\rho(t)\omega_n^2(t)}{2} \phi_{n,s}^2 \right]$$

$$= \sum_n \omega_n(t) a_{n,s}^\dagger(t) a_{n,s}(t) \quad (+\text{const})$$

$$\left( a_{n,s}(t) \equiv \sqrt{\frac{\rho(t)\omega_n(t)}{2}} \phi_{n,s} + \frac{i}{\sqrt{2\rho(t)\omega_n(t)}} \pi_{n,s} \right) \leftarrow$$

**time-dependent  
even in the  
Schr. picture**



**instantaneous ground state at  $t$  :**

$$\boxed{|0_t, t\rangle \text{ s.t. } a_{n,s}(t) |0_t, t\rangle = 0 (\forall n)}$$

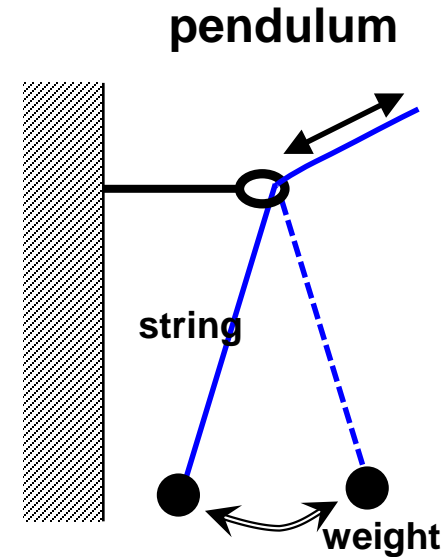
# ■ QM with time-dependent Hamiltonian

To simplify expressions, we set (for a while)

$$\phi_n \rightarrow q, \quad \pi_n \rightarrow p, \quad \omega_n(t) \rightarrow \omega(t)$$

$$H_{n,s}(t) \rightarrow H_s(t) = \frac{p_s^2}{2\rho(t)} + \frac{\rho(t)\omega^2(t)}{2} q_s^2$$

$$= \omega(t) a_s^\dagger(t) a_s(t) \quad (+\text{const})$$



Time evolution operator

$$U(t', t) = \mathcal{T} \exp \left[ -i \int_t^{t'} dt'' H_s(t'') \right]$$

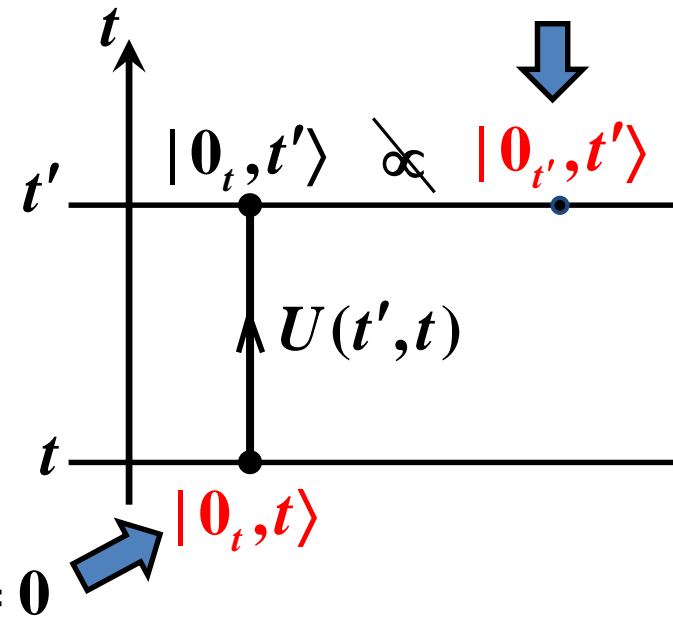
$$|\psi, t\rangle \rightarrow |\psi, t'\rangle = U(t', t) |\psi, t\rangle$$

Instantaneous ground state at time  $t$

$$|0_t, t\rangle \text{ s.t. } a_s(t) |0_t, t\rangle = 0$$

**NB:**  $|0_t, t\rangle \not\propto |0_{t'}, t'\rangle$

$$a_s(t') = 0$$





• Heisenberg picture

[Schrödinger]

$$|\psi, t\rangle = U(t, T_s) |\psi, T_s\rangle$$

$$O_s(t)$$

[Heisenberg]

$$|\psi\rangle = |\psi, T_s\rangle$$

$$O(t) = U^{-1}(t, T_s) O_s(t) U(t, T_s)$$

$\Leftrightarrow$

e.g.  $a_s(t) = \sqrt{\frac{\rho(t)\omega(t)}{2}} q_s + \frac{i}{\sqrt{2\rho(t)\omega(t)}} p_s$  ( $\langle O \rangle_t = \langle \psi, t | O_s(t) | \psi, t \rangle = \langle \psi | O(t) | \psi \rangle$ )

$\Rightarrow H(t) = \frac{1}{2\rho(t)} p^2(t) + \frac{\rho(t)\omega^2(t)}{2} q^2(t)$

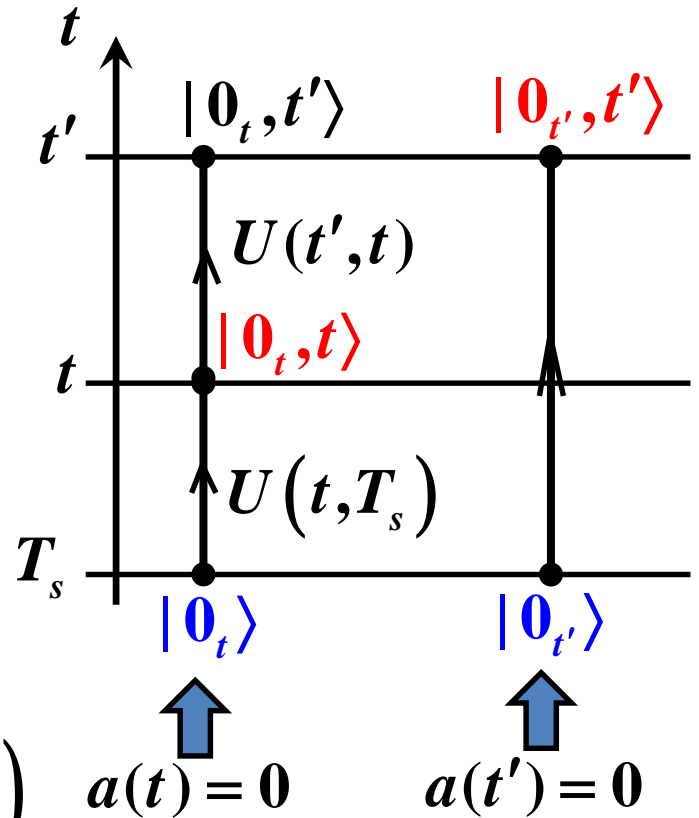
$$= \omega(t) a^\dagger(t) a(t)$$

Heisenberg eqn

$$\begin{cases} \dot{q}(t) = i[H(t), q(t)] = \frac{p(t)}{\rho(t)} \\ \dot{p}(t) = i[H(t), p(t)] = \rho(t)\omega^2(t)q(t) \end{cases}$$

$$\Leftrightarrow \ddot{q}(t) + \frac{\dot{\rho}(t)}{\rho(t)} \dot{q}(t) + \omega^2(t) q(t) = 0$$

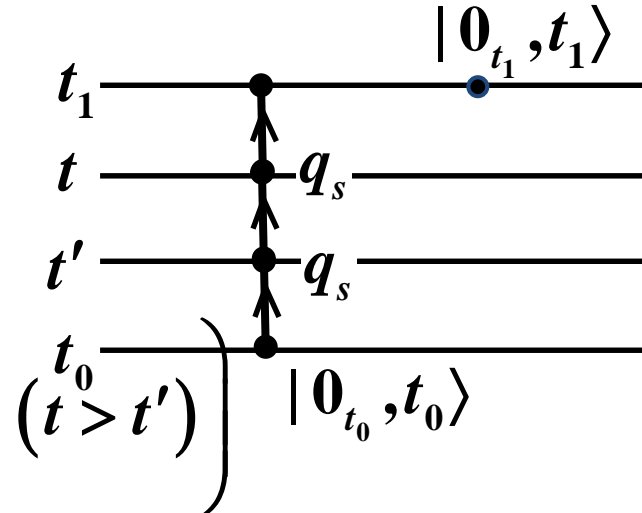
( $\Rightarrow W_\rho[f, g] \equiv \rho(f\dot{g} - \dot{f}g) : \underline{t\text{-indep}}$ )



• 2-pt function

$$G_{10}(t, t'; t_1, t_0) \equiv \frac{\langle \mathbf{0}_{t_1} | \mathbb{T} q(t) q(t') | \mathbf{0}_{t_0} \rangle}{\langle \mathbf{0}_{t_1} | \mathbf{0}_{t_0} \rangle}$$

$$\left( = \frac{\langle \mathbf{0}_{t_1}, t_1 | U(t_1, t) q_s U(t, t') q_s U(t', t_0) | \mathbf{0}_{t_0}, t_0 \rangle}{\langle \mathbf{0}_{t_1}, t_1 | U(t_1, t_0) | \mathbf{0}_{t_0}, t_0 \rangle} \right)$$



This can be easily calculated if one expresses  $q(t)$  as

$\varphi(t; t_I)$  : “wave function”

$$q(t) = \varphi(t; t_I) a(t_I) + \varphi^*(t; t_I) a^\dagger(t_I) \quad (t_I = t_0, t_1)$$

$$G_{10}(t, t'; t_1, t_0) \equiv \frac{\langle \mathbf{0}_{t_1} | a(t_1) a^\dagger(t_0) | \mathbf{0}_{t_0} \rangle}{\langle \mathbf{0}_{t_1} | \mathbf{0}_{t_0} \rangle} \varphi(t; t_1) \varphi^*(t'; t_0)$$

$$= \frac{i}{W_\rho[\varphi(s; t_1), \varphi^*(s; t_0)]} \varphi(t; t_1) \varphi^*(t'; t_0)$$

- How can one find the wave functions  $\varphi(t; t_I)$ ?

Start from Heisenberg's eqn :  $\ddot{q}(t) + \frac{\dot{\rho}(t)}{\rho(t)}\dot{q}(t) + \omega^2(t)q(t) = 0$



Find a pair of independent c-# solns:  $\{f(t), g(t)\}$



$q(t), p(t)$  then can be expressed as

$c_1, c_2$  : const operators

$$q(t) = (f(t), g(t)) \begin{pmatrix} c_1 \\ c_2 \end{pmatrix}, \quad p(t) = \rho(t)\dot{q}(t) = \rho(t) \begin{pmatrix} \dot{f}(t) & \dot{g}(t) \end{pmatrix} \begin{pmatrix} c_1 \\ c_2 \end{pmatrix}$$

$$\Leftrightarrow \begin{pmatrix} f(t) & g(t) \\ \dot{f}(t) & \dot{g}(t) \end{pmatrix} \begin{pmatrix} c_1 \\ c_2 \end{pmatrix} = \begin{pmatrix} q(t) \\ \frac{p(t)}{\rho(t)} \end{pmatrix} = \begin{pmatrix} \frac{1}{\sqrt{2\rho(t)\omega(t)}} (a(t) + a^\dagger(t)) \\ -i\sqrt{\frac{\omega(t)}{2\rho(t)}} (a(t) - a^\dagger(t)) \end{pmatrix}$$

$$= \frac{1}{\sqrt{2\rho(t)\omega(t)}} \begin{pmatrix} 1 & 1 \\ -i\omega(t) & i\omega(t) \end{pmatrix} \begin{pmatrix} a(t) \\ a^\dagger(t) \end{pmatrix}$$

$$\begin{aligned}
\therefore \begin{pmatrix} c_1 \\ c_2 \end{pmatrix} &= \frac{1}{\sqrt{2\rho(t)\omega(t)}} \begin{pmatrix} f(t) & g(t) \\ \dot{f}(t) & \dot{g}(t) \end{pmatrix}^{-1} \begin{pmatrix} 1 & 1 \\ -i\omega(t) & i\omega(t) \end{pmatrix} \begin{pmatrix} a(t) \\ a^\dagger(t) \end{pmatrix} \\
&= \frac{1}{W_\rho[f, g]} \begin{pmatrix} \mathbf{v}(t) & \bar{\mathbf{v}}(t) \\ -u(t) & -\bar{u}(t) \end{pmatrix} \begin{pmatrix} a(t) \\ a^\dagger(t) \end{pmatrix} \\
&\quad \equiv \mathbf{C}(t)
\end{aligned}$$

with

$$\begin{cases} u(t) \\ \bar{u}(t) \end{cases} \equiv \dot{f}(t) \pm i\omega(t) f(t), \quad \begin{cases} \mathbf{v}(t) \\ -\bar{\mathbf{v}}(t) \end{cases} \equiv \dot{g}(t) \pm i\omega(t) g(t)$$

$c_1, c_2$  are  $t$ -independent



$$\begin{pmatrix} c_1 \\ c_2 \end{pmatrix} = C(t) \begin{pmatrix} a(t) \\ a^\dagger(t) \end{pmatrix} = C(t') \begin{pmatrix} a(t') \\ a^\dagger(t') \end{pmatrix}$$



**Fundamental formula** [MF-Sakatani-Sugishita]

$$\begin{pmatrix} a(t) \\ a^\dagger(t) \end{pmatrix} = B(t;t') \begin{pmatrix} a(t') \\ a^\dagger(t') \end{pmatrix}$$

with

$$B(t;t') \equiv C^{-1}(t)C(t') \equiv \begin{pmatrix} \alpha^*(t;t') & -\beta^*(t;t') \\ -\beta(t;t') & \alpha(t;t') \end{pmatrix} \quad \text{“Bogoliubov coefficients”}$$

$$= \frac{-i}{2W_\rho[f, g]\sqrt{\rho(t)\omega(t)\rho(t')\omega(t')}} \begin{pmatrix} \bar{\mathbf{v}}(t)u(t') - \bar{u}(t)\mathbf{v}(t') & \bar{\mathbf{v}}(t)\bar{u}(t') - \bar{u}(t)\bar{\mathbf{v}}(t') \\ u(t)\mathbf{v}(t') - \mathbf{v}(t)u(t') & u(t)\bar{\mathbf{v}}(t') - \bar{\mathbf{v}}(t)\bar{u}(t') \end{pmatrix}$$

In particular,

$$q(t) = \frac{1}{\sqrt{2\rho(t)\omega(t)}} [a(t) + a^\dagger(t)]$$

$$= \frac{1}{\sqrt{2\rho(t)\omega(t)}} \left[ (\alpha^*(t;t_I) - \beta(t;t_I))a(t_I) + (\alpha(t;t_I) - \beta^*(t;t_I))a^\dagger(t_I) \right]$$

$$\equiv \varphi(t;t_I) a(t_I) + \varphi^*(t;t_I) a^\dagger(t_I)$$



wave functions

$$\varphi(t;t_I) = \frac{1}{W_\rho[f,g]\sqrt{2\rho_I\omega_I}} [\mathbf{v}_I f(t) - u_I g(t)]$$

$$\varphi^*(t;t_I) = \frac{1}{W_\rho[f,g]\sqrt{2\rho_I\omega_I}} [\bar{\mathbf{v}}_I f(t) - \bar{u}_I g(t)]$$

$$\left( \begin{array}{l} \rho_I \equiv \rho(t_I), \quad \omega_I \equiv \omega(t_I), \\ \left\{ \begin{array}{l} u_I \\ \bar{u}_I \end{array} \right\} \equiv \dot{f}(t_I) \pm i\omega(t_I) f(t_I), \quad \left\{ \begin{array}{l} \mathbf{v}_I \\ \bar{\mathbf{v}}_I \end{array} \right\} \equiv \dot{g}(t_I) \pm i\omega(t_I) g(t_I) \end{array} \right)$$

## • Comments

**(1) Everything is determined automatically.**

**In particular, we need not to specify the asymptotic forms of wave functions.**

**(2) Bogoliubov coefficients are independent of**

**the choice of indep solns  $\{f(t), g(t)\}$  (as they should).**

**(3) In order to ensure the convergence in the limit**  $\begin{cases} t_0 \rightarrow -\infty, \\ t_1 \rightarrow +\infty, \end{cases}$   
**we need to introduce a small imaginary part:**

$$H(t) = e^{-i\varepsilon} H_{\text{original}}(t) \Leftrightarrow \begin{cases} \rho(t) = e^{i\varepsilon} \rho_{\text{original}}(t), \\ \omega(t) = e^{-i\varepsilon} \omega_{\text{original}}(t) \end{cases}$$

# Summary

algorithm to obtain propagators [MF-Sakatani-Sugishita [1] 1301.7352]

- Decompose  $\phi(x) = \phi(t, \mathbf{x}) = \sum_n \phi_n(t) Y_n(\mathbf{x})$

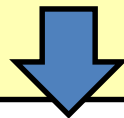
- Solve EOM to obtain c-# soln:

$$\ddot{\phi}_n(t) + \frac{\dot{\rho}(t)}{\rho(t)} \dot{\phi}_n(t) + \omega_n^2(t) \phi_n(t) = 0 \quad \longrightarrow \quad \{f_n(t), g_n(t)\}$$

- Calculate wave functions ( $t_I = t_0, t_1$ )

$$\left\{ \begin{aligned} \varphi_n(t; t_I) &= \frac{1}{W_\rho[f_n, g_n] \sqrt{2\rho(t_I)\omega_n(t_I)}} \left[ \mathbf{v}_n(t_I) f_n(t) - u_n(t_I) g_n(t) \right] \\ \varphi_n^*(t; t_I) &= \frac{1}{W_\rho[f_n, g_n] \sqrt{2\rho(t_I)\omega_n(t_I)}} \left[ \bar{\mathbf{v}}_n(t_I) f_n(t) - \bar{u}_n(t_I) g_n(t) \right] \end{aligned} \right.$$

$$\left( \begin{aligned} \left\{ \begin{aligned} u_n(t_I) \\ \bar{u}_n(t_I) \end{aligned} \right\} &\equiv \dot{f}_n(t_I) \pm i\omega_n(t_I) f_n(t_I), & \left\{ \begin{aligned} \mathbf{v}_n(t_I) \\ \bar{\mathbf{v}}_n(t_I) \end{aligned} \right\} &\equiv \dot{g}_n(t_I) \pm i\omega_n(t_I) g_n(t_I) \end{aligned} \right)$$





- **2-pt functions for each mode**

$$G_{n,10}(t, t'; t_1, t_0) = \frac{\langle \mathbf{0}_{t_1} | \mathbf{T} \phi_n(t) \phi_n(t') | \mathbf{0}_{t_0} \rangle}{\langle \mathbf{0}_{t_1} | \mathbf{0}_{t_0} \rangle} = \frac{i \phi_n(t; t_1) \phi_n^*(t'; t_0)}{W_\rho[\phi_n(s; t_1), \phi_n^*(s; t_0)]}$$

$$G_{n,00}(t, t'; t_0, t_0) = \frac{\langle \mathbf{0}_{t_0} | \mathbf{T} \phi_n(t) \phi_n(t') | \mathbf{0}_{t_0} \rangle}{\langle \mathbf{0}_{t_0} | \mathbf{0}_{t_0} \rangle} = \frac{i \phi_n(t; t_0) \phi_n^*(t'; t_0)}{W_\rho[\phi_n(s; t_0), \phi_n^*(s; t_0)]}$$

- **Sum up over  $n$ :**

$$G_{10}(x, x'; t_1, t_0) = \sum_n G_{n,10}(t, t'; t_1, t_0) Y_n(\mathbf{x}) Y_n(\mathbf{x}')$$

$$G_{00}(x, x'; t_0, t_0) = \sum_n G_{n,00}(t, t'; t_0, t_0) Y_n(\mathbf{x}) Y_n(\mathbf{x}')$$

- **Take the limit  $t_0 \rightarrow -\infty, t_1 \rightarrow +\infty$ ,**

**to obtain the in-out and in-in propagators:**

$$G^{\text{out/in}}(x, x') \equiv \lim_{\substack{t_0 \rightarrow -\infty \\ t_1 \rightarrow +\infty}} G_{10}(x, x'; t_1, t_0)$$

$$G^{\text{in/in}}(x, x') \equiv \lim_{t_0 \rightarrow -\infty} G_{00}(x, x'; t_0, t_0)$$

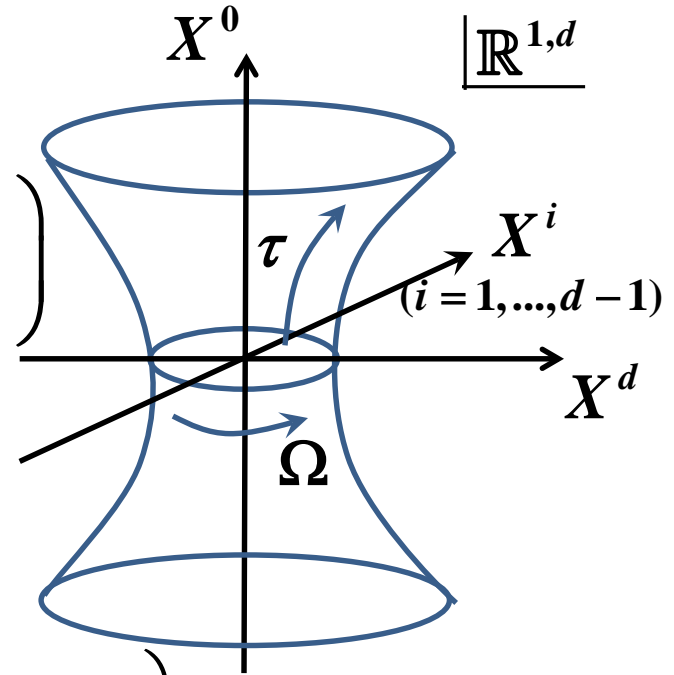
### **3. Propagators in de Sitter space**

**[1] MF-Sakatani-Sugishita [PRD88 (2013) 024041, 1301.7352]**

## Global patch

$$ds^2 = -d\tau^2 + \cosh^2 \tau d\Omega_{d-1}^2 \quad (\ell = 1)$$

$$= -\frac{dt^2}{(1-t^2)^2} + \frac{1}{1-t^2} d\Omega_{d-1}^2 \quad \left( \begin{array}{l} t \equiv \tanh \tau; \\ -1 < t < +1 \end{array} \right)$$



## mode expansion

$$\phi(x) = \sum_{L \geq 0} \sum_{M=1}^{N_L^{(d)}} \phi_{LM}(t) Y_{LM}(\Omega)$$

$$\left( \begin{array}{l} \Delta_{d-1} Y_{LM}(\Omega) = -L(L+d-2) Y_{LM}(\Omega) \\ Y_{LM}(\Omega) \in \mathbb{R}, \quad N_L^{(d)} = \frac{(L+d-2)!}{(d-2)!L!} (2L+d-2) \\ \int d^{d-1}\Omega Y_{LM}(\Omega) Y_{L'M'}(\Omega) = \delta_{LL'} \delta_{MM'} \end{array} \right)$$

## EOM

$$\ddot{\phi}_{LM}(t) + \frac{(d-3)t}{1-t^2} \dot{\phi}_{LM}(t) + \left[ \frac{L(L+d-2)}{1-t^2} + \frac{m^2}{(1-t^2)^2} \right] \phi_{LM}(t) = 0$$



• independent c-# solns

$$\begin{cases} f_{LM}(t) = (1-t^2)^{(d-1)/4} \mathbf{P}_{L+(d-3)/2}^\nu(t) \\ g_{LM}(t) = (1-t^2)^{(d-1)/4} \mathbf{Q}_{L+(d-3)/2}^\nu(t) \end{cases}$$

$$\nu \equiv \begin{cases} \sqrt{\left(\frac{d-1}{2}\right)^2 - m^2} & \left(m < \frac{d-1}{2}\right) \\ i\sqrt{m^2 - \left(\frac{d-1}{2}\right)^2} & \left(m > \frac{d-1}{2}\right) \end{cases} \equiv i\mu$$

• wave functions

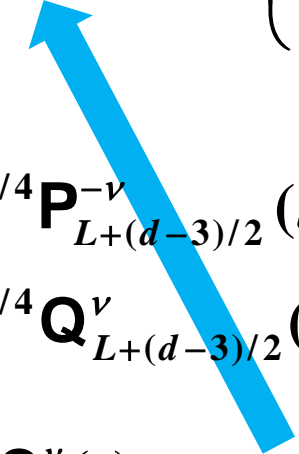
$$\begin{cases} \varphi_{LM}(t; t_1) \propto (1-t^2)^{(d-1)/4} \mathbf{P}_{L+(d-3)/2}^{-\nu}(t) & (t_1 \sim +1) \\ \varphi_{LM}(t; t_0) \propto (1-t^2)^{(d-1)/4} \mathbf{Q}_{L+(d-3)/2}^\nu(t) & (t_0 \sim -1) \end{cases}$$

• 2-pt functions

$\left( \mathbf{P}_k^\nu(t), \mathbf{Q}_k^\nu(t) : \text{associated Legendre functions} \right)$   
 defined on the interval  $-1 < t < +1$

$$G_{LM}^{\text{out/in}}(t, t')$$

$$= \begin{cases} \frac{i\pi}{2 \sin \pi \nu} \left[ (1-t_>^2)(1-t_<^2) \right]^{(d-1)/4} \mathbf{P}_{L+(d-3)/2}^{-\nu}(t_>) \mathbf{P}_{L+(d-3)/2}^\nu(t_>) & (d : \text{odd}) \\ \frac{1}{\cos \pi \nu} \left[ (1-t_>^2)(1-t_<^2) \right]^{(d-1)/4} \mathbf{P}_{L+(d-3)/2}^{-\nu}(t_>) \mathbf{Q}_{L+(d-3)/2}^\nu(t_>) & (d : \text{even}) \end{cases}$$



• mode sum

$$G^{\text{out/in}}(x, x')$$

$$= \sum_L \sum_M G_{LM}^{\text{out/in}}(t, t') Y_{LM}(\Omega) Y_{LM}(\Omega') = \sum_L G_L^{\text{out/in}}(t, t') \underbrace{\sum_M Y_{LM}(\Omega) Y_{LM}(\Omega')}_{\frac{2L+d-2}{(d-2)|\Omega_{d-1}|} C_L^{(d-2)/2}(\Omega \cdot \Omega') \text{ Gegenbauer polynomial}}$$

$$= \begin{cases} \frac{-e^{-i\pi(d-2)/2}}{2(2\pi)^{d/2} \sin \pi\nu} \begin{bmatrix} (u_+^2 - 1)^{-(d-2)/4} Q_{\nu-1/2}^{(d-2)/2}(u_+) \\ -(u_-^2 - 1)^{-(d-2)/4} Q_{\nu-1/2}^{(d-2)/2}(u_-) \end{bmatrix} & (d : \text{odd}) \\ \frac{ie^{-i\pi(d-2)/2}}{2(2\pi)^{d/2} \cos \pi\nu} \begin{bmatrix} (u_+^2 - 1)^{-(d-2)/4} Q_{\nu-1/2}^{(d-2)/2}(u_+) \\ +(u_-^2 - 1)^{-(d-2)/4} Q_{\nu-1/2}^{(d-2)/2}(u_-) \end{bmatrix} & (d : \text{even}) \end{cases}$$

$\left( P_k^\nu(z), Q_k^\nu(z) : \text{associated Legendre functions} \right)$   
 defined except for the cut  $-\infty < z < +1$

Here  $u_\pm(x, x') = -Z(x, x') \pm i0$ .

  $G^{\text{out/in}}(x, x')$  is de Sitter invariant!

Similarly,

$$G_{\text{odd}}^{\text{in/in}}(x, x') = G_{\text{odd}}^{\text{out/in}}(x, x')$$

$$= \frac{-e^{-i\pi(d-2)/2}}{2(2\pi)^{d/2} \sin \pi\nu} \left[ \begin{array}{c} (u_+^2 - 1)^{-(d-2)/4} Q_{\nu-1/2}^{(d-2)/2}(u_+) \\ -(u_-^2 - 1)^{-(d-2)/4} Q_{\nu-1/2}^{(d-2)/2}(u_-) \end{array} \right]$$

$$G_{\text{even}}^{\text{in/in}}(x, x') = \frac{e^{-i\pi(d-2)/2}}{2(2\pi)^{d/2} \sin \pi\nu} \left\{ \begin{array}{l} \left[ \begin{array}{c} (u_+^2 - 1)^{-(d-2)/4} Q_{\nu-1/2}^{(d-2)/2}(u_+) \\ -(u_-^2 - 1)^{-(d-2)/4} Q_{\nu-1/2}^{(d-2)/2}(u_-) \end{array} \right] \\ + \frac{i\pi}{\cos \pi\nu} \left[ \begin{array}{c} e^{-i\pi\nu} (u_+^2 - 1)^{-(d-2)/4} P_{\nu-1/2}^{(d-2)/2}(u_+) \\ + e^{i\pi\nu} (u_-^2 - 1)^{-(d-2)/4} P_{\nu-1/2}^{(d-2)/2}(u_-) \end{array} \right] \end{array} \right\}$$

Again,  $u_{\pm}(x, x') = -Z(x, x') \pm i0$

  $G^{\text{in/in}}(x, x')$  is also de Sitter invariant!

# ■ Poincare patch

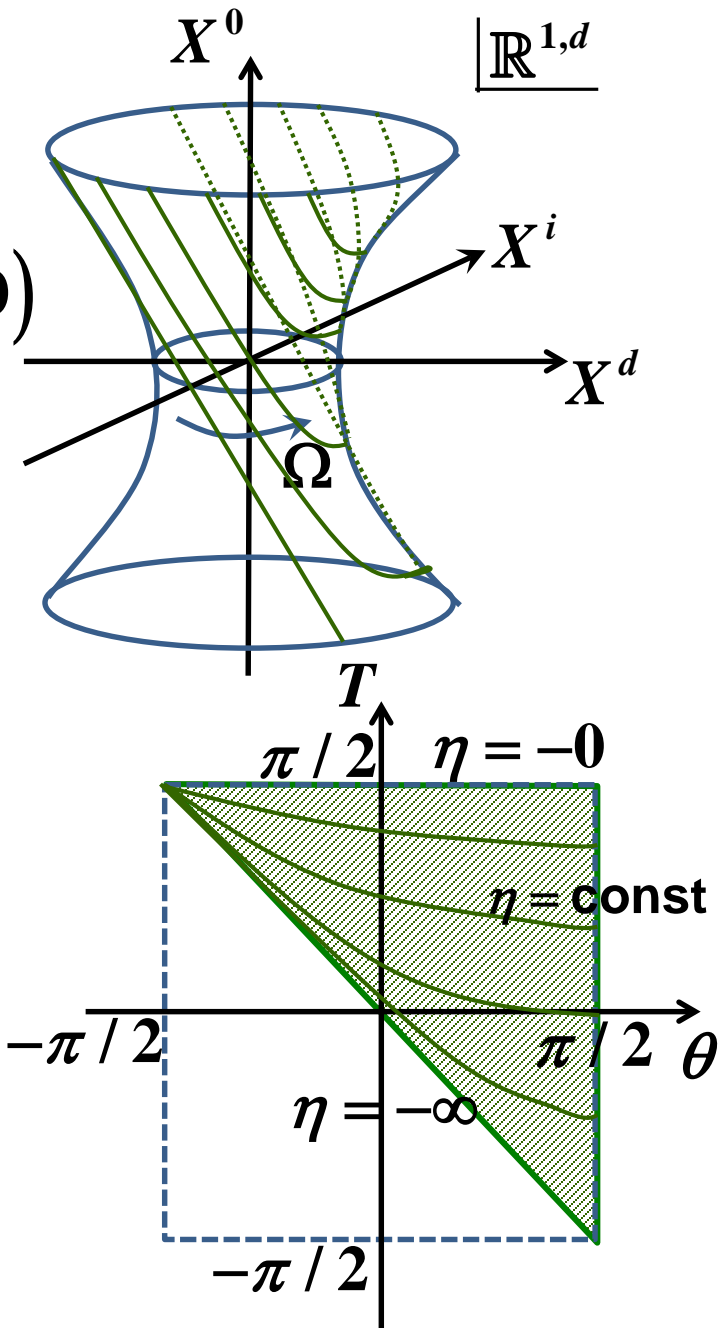
$$ds^2 = -d\tau^2 + e^{2\tau} dx^2 \quad (\ell = 1)$$

$$= \frac{-d\eta^2 + dx^2}{\eta^2} \quad (\eta \equiv -e^{-\tau}; -\infty < \eta < 0)$$

## ● mode expansion

$$\phi(x) = \sum_{\mathbf{k} \geq 0} \sum_a \phi_{\mathbf{k},a}(t) Y_{\mathbf{k},a}(\mathbf{x})$$

$$\left( \begin{array}{l} \Delta_{d-1} Y_{\mathbf{k},a}(\mathbf{x}) = -k^2 Y_{\mathbf{k},a}(\mathbf{x}) \quad (k = |\mathbf{k}|) \\ \left\{ \begin{array}{l} Y_{\mathbf{k}=0, a=1} = \frac{1}{\sqrt{V}} \quad \left( \mathbf{k} = (k_1, \dots, k_{d-1}) > 0 \right) \\ \quad \quad \quad \quad \quad \quad \quad (\Leftrightarrow \text{1st nonzero element} > 0) \\ Y_{\mathbf{k}>0, a=1} = \sqrt{\frac{2}{V}} \cos \mathbf{k} \cdot \mathbf{x}, \quad Y_{\mathbf{k}>0, a=2} = \sqrt{\frac{2}{V}} \sin \mathbf{k} \cdot \mathbf{x} \end{array} \right. \\ \int_V d^{d-1} \mathbf{x} Y_{\mathbf{k},a}(\mathbf{x}) Y_{\mathbf{k}',a'}(\mathbf{x}) = \delta_{\mathbf{k},\mathbf{k}'} \delta_{a,a'} \end{array} \right)$$



- EOM

$$\ddot{\phi}_{\mathbf{k},a}(\eta) - \frac{d-2}{\eta} \dot{\phi}_{\mathbf{k},a}(\eta) + \left( k^2 + \frac{m^2}{\eta^2} \right) \phi_{\mathbf{k},a}(\eta) = 0$$

- independent c-# solns

$$\begin{cases} f_{LM}(\eta) = (-\eta)^{(d-1)/2} J_\nu(-k\eta) \\ g_{LM}(\eta) = (-\eta)^{(d-1)/2} N_\nu(-k\eta) \end{cases} \quad \nu \equiv \begin{cases} \sqrt{\left(\frac{d-1}{2}\right)^2 - m^2} & \left( m < \left(\frac{d-1}{2}\right) \right) \\ i\sqrt{m^2 - \left(\frac{d-1}{2}\right)^2} = i\mu & \left( m > \left(\frac{d-1}{2}\right) \right) \end{cases}$$

- 2-pt functions

$$G_{\mathbf{k},a}^{\text{out/in}}(\eta, \eta') = \frac{\pi}{2} [(-\eta)(-\eta')]^{(d-1)/2} J_\nu(-k\eta_>) H_\nu^{(2)}(-k\eta_<)$$

$$G_{\mathbf{k},a}^{\text{in/in}}(\eta, \eta') = \frac{\pi}{4} [(-\eta)(-\eta')]^{(d-1)/2} H_\nu^{(1)}(-k\eta_<) H_\nu^{(2)}(-k\eta_<)$$





- mode sum

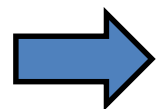
$$G^{\text{out/in}}(x, x') = \frac{e^{-i\pi(d-2)}}{(2\pi)^{d/2}} (u^2 - 1)^{-(d-2)/4} Q_{\nu-1/2}^{(d-2)/2}(u)$$

$$G^{\text{in/in}}(x, x')$$

$$= \frac{\Gamma\left(\frac{d-1}{2} + \nu\right) \Gamma\left(\frac{d-1}{2} - \nu\right)}{2(2\pi)^{d/2}} (u^2 - 1)^{-(d-2)/4} P_{\nu-1/2}^{-(d-2)/2}(u)$$

$$= \frac{\Gamma\left(\frac{d-1}{2} + \nu\right) \Gamma\left(\frac{d-1}{2} - \nu\right)}{(4\pi)^{d/2} \Gamma(d/2)} F\left(\frac{d-1}{2} + \nu, \frac{d-1}{2} - \nu, \frac{d}{2}, \frac{1-u}{2}\right)$$

Here,  $u = -Z(x, x') + i0$ .



$G^{\text{out/in}}(x, x')$ ,  $G^{\text{in/in}}(x, x')$  are de Sitter invariant!

## ■ Properties

### • in-in propagators in the Poincare patch

- no massless limit (as is already known)
- agrees with the propagators for the BD vacuum:

$$G^{\text{in/in}}(x, x') = G_{\text{BD}}(x, x') \quad \text{i.e.} \quad |0_{\eta_0}\rangle \xrightarrow{\eta_0 \rightarrow -\infty} |\text{BD}\rangle$$

➔  $G^{\text{in/in}}(x, x')$  : analytic continuation of

Euclidean propagator on  $S^d$

$$\left( ds^2 = -d\tau^2 + \cosh^2 \tau d\Omega_{d-1}^2 \iff ds_E^2 = d\tau_E^2 + \cos^2 \tau_E d\Omega_{d-1}^2 \right)$$

### • in-out propagators in the Poincare patch

- can take the massless limit
- $G^{\text{out/in}}(x, x')$  : analytic continuation of

propagator on Euclidean AdS

$$\left( ds^2 = \ell^2 \frac{-d\eta^2 + d\mathbf{x}^2}{\eta^2} \iff ds_E^2 = \ell_E^2 \frac{d\eta_E^2 + d\mathbf{x}^2}{\eta_E^2} \quad (\ell_E = i\ell, \eta_E = i\eta) \right)$$

## ■ in-in and in-out

Which propagator should be used depends on the problem under consideration.

- in-in 
$$G^{\text{in/in}}(x, x') = \frac{\langle \text{in} | \mathbf{T} \phi(x) \phi(x') | \text{in} \rangle}{\langle \text{in} | \text{in} \rangle}$$

In particular,

$$\begin{aligned} & G^{\text{in/in}}(t, \mathbf{x}; t, \mathbf{x}') \\ &= \left( \phi_s(\mathbf{x}) U(t, -\infty) | \mathbf{0}_{-\infty}, -\infty \rangle \right)^\dagger \phi_s(\mathbf{x}') U(t, -\infty) | \mathbf{0}_{-\infty}, -\infty \rangle \end{aligned}$$



suitable when discussing expectation values and quantum fluctuations

cf. quantum fluct. in inflation

- in-out 
$$G^{\text{out/in}}(x, x') = \frac{\langle \text{out} | \mathbf{T} \phi(x) \phi(x') | \text{in} \rangle}{\langle \text{out} | \text{in} \rangle}$$



suitable when discussing the S-matrix

## ■ More on the in-out propagators

- relation to the functional integral with  $i\varepsilon$

$$G^{\text{out/in}}(x, x') = \frac{\int [d\phi(x)] e^{iS_\varepsilon[\phi]} \phi(x) \phi(x')}{\int [d\phi(x)] e^{iS_\varepsilon[\phi]}}$$

← (proved and confirmed numerically)

with  $S_\varepsilon[\phi] = \int d^d x \sqrt{-g} \left[ -\frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - \frac{m^2 - i\varepsilon}{2} \phi^2 \right]$

- relation to the random walks in de Sitter space

∃ a heat kernel representation:

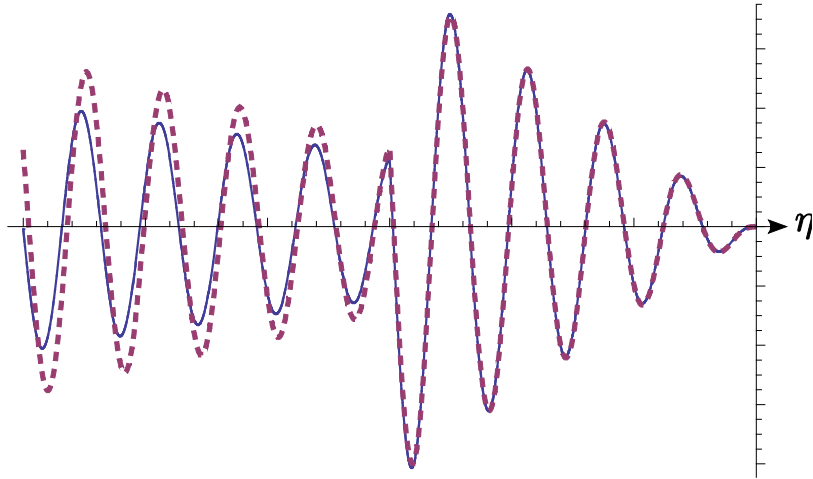
$$G^{\text{out/in}}(x, x') = \frac{1}{2} \int_0^\infty dT \langle x | e^{-i(T/2)(-\square + m^2 - i\varepsilon)} | x' \rangle = \sum_{\substack{\text{path} \\ \text{from } x' \text{ to } x}} e^{i \text{length}(\text{path})}$$

( =  $\langle x | \frac{-i}{-\square + m^2} | x' \rangle$  )

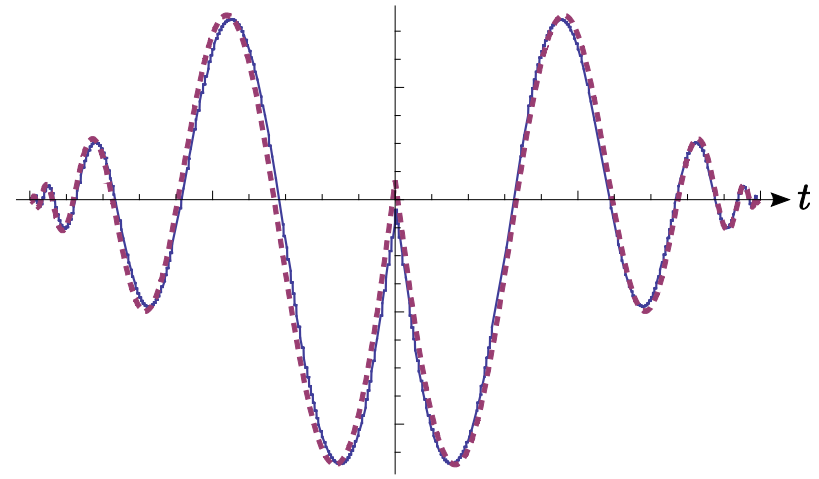
→  $\frac{\partial}{\partial m^2} G^{\text{out/in}}(x, x') = i \int d^d y \sqrt{-g(y)} G^{\text{out/in}}(x, y) G^{\text{out/in}}(y, x')$

“composition law” [Polyakov]

## ■ Numerical results



$G_k^{\text{out/in}}(\eta, 0)$  in Poincare  
( $d = 4, m \ell = 0.5, k = 1$ )



$G_L^{\text{out/in}}(t, 0)$  in global  
( $d = 4, m \ell = 9, L = 1$ )

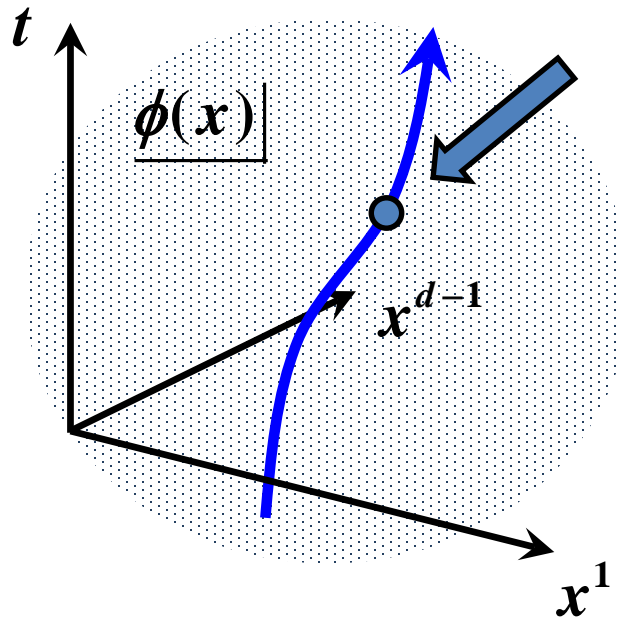
## **PART 2 : Unruh-DeWitt detector in curved spacetime**

**[2] MF-Sakatani-Sugishita [PRD89 (2014) 064024, 1305.0256]**

## **4. Motivation and the result**

**[2] MF-Sakatani-Sugishita [PRD89 (2014) 064024, 1305.0256]**

## ■ Unruh-DeWitt detector



detector

{
   
 interacting with  $\phi(x)$ 
  
 moving along a classical path:
   
 $x(\tau) = (t(\tau), \mathbf{x}(\tau)) = (t, \mathbf{x}(t))$

detector Hamiltonian  $H^d$

{
   
 const for proper time  $\tau$ 
  
 $H^d |n\rangle = E_n |n\rangle \quad (n = 0, 1, 2, \dots)$

For some spacetime,

the detector behaves as if it is in a heat bath with temp  $1/\beta$ .

$\Leftrightarrow$  The density distribution of the detector behaves as

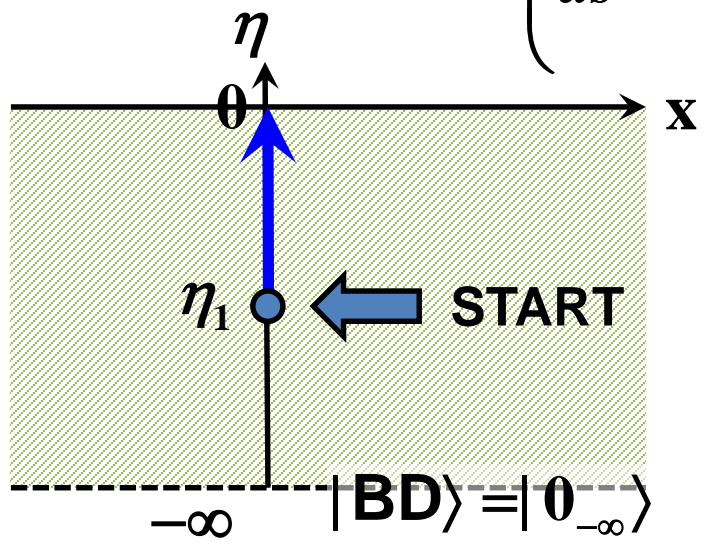
$$\rho_n(t) \xrightarrow[t \rightarrow \infty]{} \frac{1}{Z} e^{-\beta E_n} \left( Z = \sum_{k \geq 0} e^{-\beta E_k} \right)$$

This transition is instantaneous  
for an "ideal" detector.



■ de Sitter case

$$\left( ds^2 = \ell^2 \frac{-d\eta^2 + dx^2}{\eta^2} \right)$$



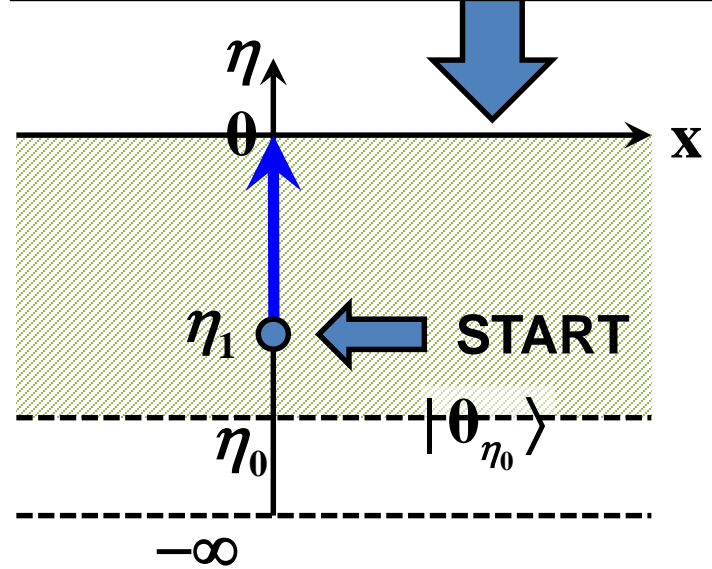
$$\rho_n(\eta) \xrightarrow[\text{relax.}]{\eta \gg \eta_1} \frac{1}{Z} e^{-2\pi\ell E_n} \quad \left( \text{Gibbs distribution} \right)$$

Detector feels as if

it is in a thermal bath with  $T = 1 / 2\pi\ell$

(Relaxation is instantaneous if the detector is "ideal".)

What if taking the instantaneous vacuum at a finite past?



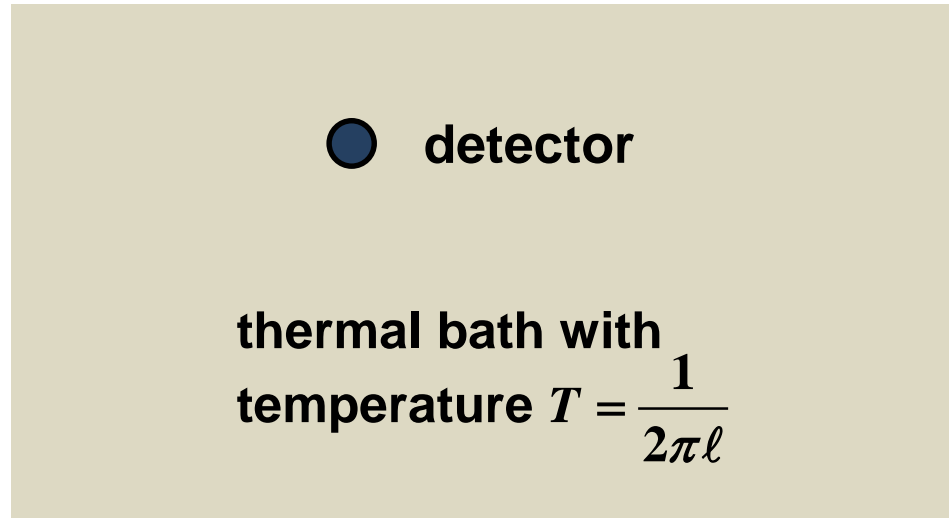
Detector will feel as if

it is in a medium which

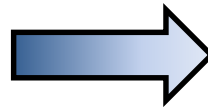
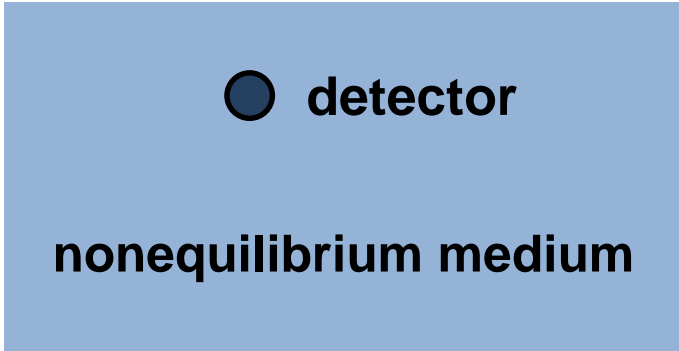
{ is not in thermal equilibrium  
relaxes to the equilibrium corr. to |BD⟩

(The difference between |0\_{η₀}⟩ and |BD⟩ becomes irrelevant as time goes on.)

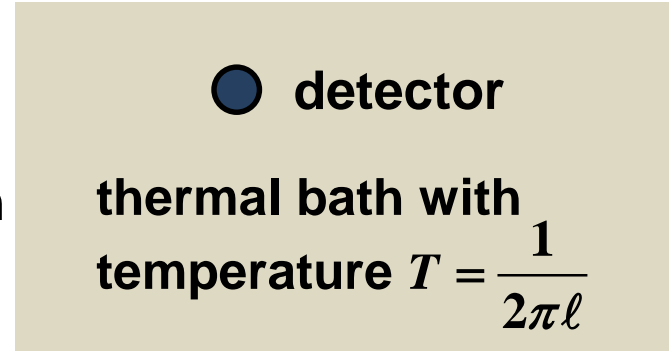
For  $|0_{-\infty}\rangle = |\text{BD}\rangle$ :



For  $|0_{\eta_0}\rangle$ :



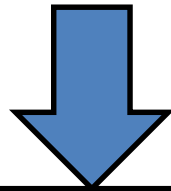
relaxation of medium  
with universal  
relaxation time  $\ell/2$



## ■ Result

**We find that**

**even for an ideal detector,  
the density distribution  $\{\rho_n(\eta)\}$  has  
a finite relaxation time  $\tau_{\text{relax}} = \ell / 2$ .**



$$\tau_{\text{relax}} = \frac{\ell}{2} : \left\{ \begin{array}{l} \text{universal quantity} \\ \text{related to nonequilibrium dynamics} \\ \text{intrinsic to de Sitter space} \end{array} \right.$$

**To show this, we first develop a framework...**

## **5. Master equation for Unruh-DeWitt detector (general theory)**

**[2] MF-Sakatani-Sugishita [PRD89 (2014) 064024, 1305.0256]**

## ■ Setup

- total Hilbert space:  $\mathcal{H}^{\text{tot}} = \mathcal{H}^{\text{d}} \otimes \mathcal{H}^{\phi}$

- Hamiltonian (Schrödinger picture)

$$H^{\text{tot}}(t) = H^{\text{d}} \frac{d\tau}{dt} \otimes 1 + 1 \otimes H^{\phi}(t) + V(t)$$

$$V(t) = \mu \frac{d\tau}{dt} \otimes \phi(\mathbf{x}(t)) \cdot \theta(t - t_1)$$

- time evolution operator

$$U^{\text{tot}}(t, t_1) = \mathbb{T} \exp \left[ -i \int_{t_1}^t dt' H^{\text{tot}}(t') \right]$$

- total density matrix

$$\rho^{\text{tot}}(t) = U^{\text{tot}}(t, t_1) \rho^{\text{tot}}(t_1) [U^{\text{tot}}(t, t_1)]^{-1}$$

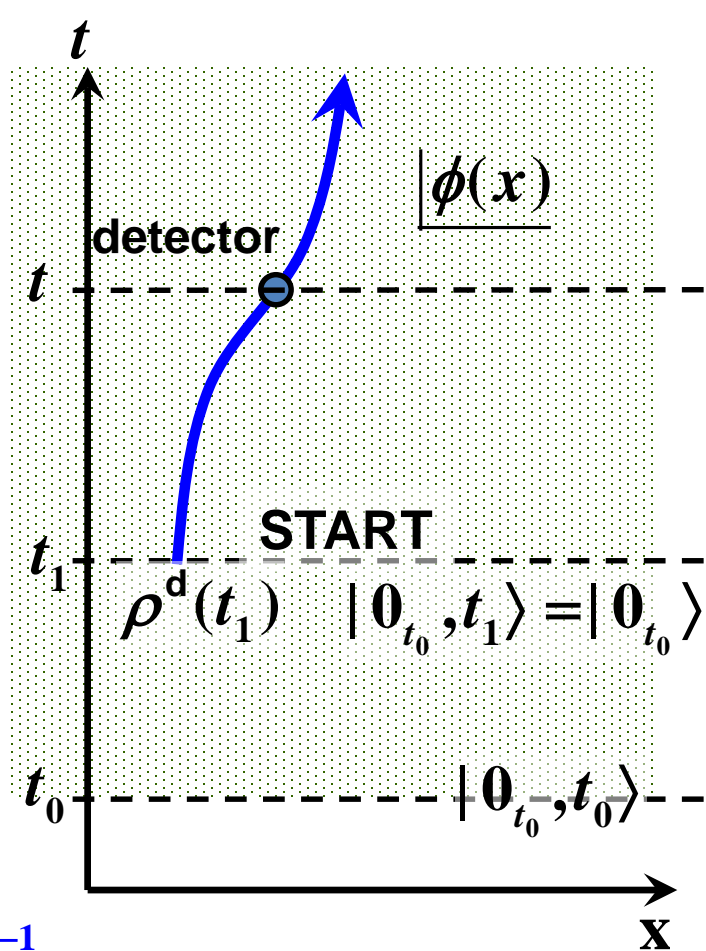
$$\Leftrightarrow \dot{\rho}^{\text{tot}}(t) = -i [H^{\text{tot}}(t), \rho^{\text{tot}}(t)] \quad \text{“von Neumann eqn”}$$

- BC

$$\rho^{\text{tot}}(t_1) = \rho^{\text{d}}(t_1) \otimes \rho^{\phi}(t_1)$$

$$\left( \text{e.g. } \rho^{\phi}(t_1) = |\mathbf{0}_{t_0}, t_1\rangle \langle \mathbf{0}_{t_0}, t_1| = |\mathbf{0}_{t_0}\rangle \langle \mathbf{0}_{t_0}| \right) \leftarrow T_s = t_1$$

detector  $\phi(x)$



Suppose that the behavior of the detector is only concerned.



- reduced density matrix

$$\rho(t) \equiv \text{Tr}_\phi \rho^{\text{tot}}(t) \quad \left( \begin{array}{l} \text{NB: } \rho(t_1) = \text{Tr}_\phi \rho^{\text{tot}}(t_1) \\ = \text{Tr}_\phi (\rho^{\text{d}}(t_1) \otimes \rho^\phi(t_1)) = \rho^{\text{d}}(t_1) \end{array} \right)$$

- interaction picture

$$\rho_I^{\text{tot}}(t) \equiv [U_0^{\text{tot}}(t, t_1)]^{-1} \rho^{\text{tot}}(t) U_0^{\text{tot}}(t, t_1)$$

$$\left( U_0^{\text{tot}}(t, t_1) = \mathbf{T}e^{-i \int_{t_1}^t dt' H_0^{\text{tot}}(t')} = e^{-i H^{\text{d}} \cdot (\tau - \tau_1)} \otimes \mathbf{T}e^{-i \int_{t_1}^t dt' H^\phi(t')} \right)$$

- von Neumann equation

$$\dot{\rho}_I^{\text{tot}}(t) = -i [V_I(t), \rho_I^{\text{tot}}(t)] = -i \text{ad}_{V_I(t)} \rho_I^{\text{tot}}(t)$$

$$\left( \begin{array}{l} V_I(t) \equiv [U_0^{\text{tot}}(t, t_1)]^{-1} V(t) U_0^{\text{tot}}(t, t_1) \\ = \mu_I(t) \frac{d\tau}{dt} \otimes \phi_I(t, \mathbf{x}(t)) \cdot \theta(t - t_1) \end{array} \right)$$

- BC

$$\rho_I^{\text{tot}}(t_1) = \rho^{\text{tot}}(t_1) = \rho^{\text{d}}(t_1) \otimes \rho^\phi(t_1)$$

■ Method of orthogonal projection [cf. Kubo-Toda-Hashitsume]

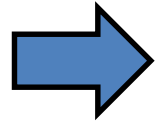
• For  $O \in \text{End } \mathcal{H}^{\text{tot}}$ , we introduce

$$\mathcal{P} : \text{End } \mathcal{H}^{\text{tot}} \rightarrow \text{End } \mathcal{H}^{\text{tot}}$$

$$O \mapsto \mathcal{P}O \equiv (\text{Tr}_\phi O) \otimes X^\phi$$

$X^\phi$  : arbitrary as long as  
it is  $\left( \begin{array}{l} \text{time indep} \\ \text{satisfying } \text{Tr}_\phi X^\phi = 1 \end{array} \right)$

$\Rightarrow \mathcal{P}^2 = \mathcal{P}$



$$\mathcal{P}\rho_I^{\text{tot}}(t) \equiv (\text{Tr}_\phi \rho_I^{\text{tot}}(t)) \otimes X^\phi = \rho_I(t) \otimes X^\phi$$

• Choose  $X^\phi$  s.t.

$$\begin{cases} X^\phi = \rho_I^\phi(t_1) = |\mathbf{0}_{t_0}\rangle\langle\mathbf{0}_{t_0}| \\ \langle\phi_I(x)\rangle_X = \text{Tr}_\phi(\phi_I(x)X^\phi) = \langle\mathbf{0}_{t_0}|\phi_I(x)|\mathbf{0}_{t_0}\rangle \equiv 0 \end{cases}$$



master equation [MF-Sakatani-Sugishita [2]] ( $\mathcal{Q} \equiv 1 - \mathcal{P}$ )

$$\frac{d\rho_I(t)}{dt} \otimes \rho_I^\phi(t_1) = -\mathcal{P} \text{ad}_{V_I(t)} \int_{t_1}^t dt' \text{T}e^{-i\int_{t'}^t dt'' \mathcal{Q} \text{ad}_{V_I(t'')}} \text{ad}_{V_I(t')} \rho_I(t') \otimes \rho_I^\phi(t_1)$$

proof Using  $P + \mathcal{Q} \equiv 1$ ,

$$\frac{d}{dt} \rho_I^{\text{tot}}(t) = -i \mathbf{ad}_{V_I(t)} \rho_I^{\text{tot}}(t) \Rightarrow \begin{cases} \frac{d}{dt} \mathcal{P} \rho_I^{\text{tot}}(t) = -i \mathcal{P} \mathbf{ad}_{V_I(t)} \mathcal{P} \rho_I^{\text{tot}}(t) - i \mathcal{P} \mathbf{ad}_{V_I(t)} \mathcal{Q} \rho_I^{\text{tot}}(t) \quad \text{--- (1)} \\ \frac{d}{dt} \mathcal{Q} \rho_I^{\text{tot}}(t) = -i \mathcal{Q} \mathbf{ad}_{V_I(t)} \mathcal{P} \rho_I^{\text{tot}}(t) - i \mathcal{Q} \mathbf{ad}_{V_I(t)} \mathcal{Q} \rho_I^{\text{tot}}(t) \quad \text{--- (2)} \end{cases}$$

This is a coupled diff eqn for  $\mathcal{P} \rho_I^{\text{tot}}(t)$  and  $\mathcal{Q} \rho_I^{\text{tot}}(t)$ .

First solve (2) w.r.t.  $\mathcal{Q} \rho_I^{\text{tot}}(t)$ :

$$\mathcal{Q} \rho_I^{\text{tot}}(t) = \mathbf{T}e^{-i \int_{t_1}^t dt' \mathcal{Q} \mathbf{ad}_{V_I(t')}} \mathcal{Q} \rho_I^{\text{tot}}(t_1) - i \int_{t_1}^t dt' \mathbf{T}e^{-i \int_{t'}^t dt'' \mathcal{Q} \mathbf{ad}_{V_I(t'')}} \mathcal{Q} \mathbf{ad}_{V_I(t')} \mathcal{P} \rho_I^{\text{tot}}(t')$$

Substitute this to (1) using  $\mathcal{P} \rho_I^{\text{tot}}(t) = \rho_I(t) \otimes X^\phi = \rho_I(t) \otimes \rho_I^\phi(t_1)$ :

$$\begin{aligned} \frac{d \rho_I(t)}{dt} \otimes \rho_I^\phi(t_1) = & \underbrace{-i \mathcal{P} \mathbf{ad}_{V_I(t)} \mathcal{P} \rho_I^{\text{tot}}(t)}_{\substack{\uparrow \\ \text{--- (*)}}} - i \mathcal{P} \mathbf{ad}_{V_I(t)} \mathbf{T}e^{-i \int_{t_1}^t dt' \mathcal{Q} \mathbf{ad}_{V_I(t')}} \underbrace{\mathcal{Q} \rho_I^{\text{tot}}(t_1)}_{\substack{\uparrow \\ \text{--- (*)}}} \\ & - \mathcal{P} \mathbf{ad}_{V_I(t)} \int_{t_1}^t dt' \mathbf{T}e^{-i \int_{t'}^t dt'' \mathcal{Q} \mathbf{ad}_{V_I(t'')}} \mathcal{Q} \mathbf{ad}_{V_I(t')} \mathcal{P} \rho_I^{\text{tot}}(t') \quad \text{--- (*)} \end{aligned}$$

We have chosen  $X^\phi = \rho_I^\phi(t_1)$  and  $\rho_I^{\text{tot}}(t_1) = \rho_I^{\text{d}}(t_1) \otimes \rho_I^\phi(t_1)$

$$\Rightarrow \mathcal{P} \rho_I^{\text{tot}}(t_1) = \rho_I(t_1) \otimes X^\phi = \rho_I^{\text{d}}(t_1) \otimes \rho_I^\phi(t_1) = \rho_I^{\text{tot}}(t_1) \Leftrightarrow \mathcal{Q} \rho_I^{\text{tot}}(t_1) = 0$$

We have set  $\text{Tr}_\phi(\phi_I(x) X^\phi) = \text{Tr}_\phi(\phi_I(x) \rho_I^\phi(t_1)) = 0$

$$\Rightarrow \mathcal{P} \mathbf{ad}_{V_I(t)} \mathcal{P} \rho_I^{\text{tot}}(t) = \text{Tr}_\phi(\phi_I(x) \rho_I^\phi(t_1)) \otimes \rho_I^\phi(t_1) = 0$$

Only the third term in (\*) survives. ■



## ■ Approximation of master equation

master equation [MF-Sakatani-Sugishita [2]] ( $\mathcal{Q} \equiv 1 - \mathcal{P}$ )

$$\frac{d\rho_I(t)}{dt} \otimes \rho_I^\phi(t_1) = -\mathcal{P} \text{ad}_{V_I(t)} \int_{t_1}^t dt' \text{Tr} e^{-i \int_{t'}^t dt'' \mathcal{Q} \text{ad}_{V_I(t'')}} \text{ad}_{V_I(t')} \rho_I(t') \otimes \rho_I^\phi(t_1)$$

• perturbation to  $O(\mu^2)$

➔ 
$$\frac{d\rho_I(t)}{dt} = -\text{Tr}_\phi \left[ \text{ad}_{V_I(t)} \int_{t_1}^t dt' \text{ad}_{V_I(t')} \rho_I(t') \otimes \rho_I^\phi(t_1) \right] + O(\mu^3)$$

With 
$$\begin{cases} V_I(t) = \mu_I(t) \frac{d\tau}{dt} \otimes \phi_I(x(\tau)) \cdot \theta(t - t_1) \\ X^\phi = \rho_I^\phi(t_1) = |\mathbf{0}_{t_0}\rangle \langle \mathbf{0}_{t_0}| \end{cases}$$

$$\begin{aligned} G_X^+(x, x') &\equiv \text{Tr}_\phi \left( \phi_I(x) \phi_I(x') X^\phi \right) \\ &= \langle \mathbf{0}_{t_0} | \phi_I(x) \phi_I(x') | \mathbf{0}_{t_0} \rangle \end{aligned}$$

we have

$$\dot{\rho}_I(\tau) = \int_{\tau_1}^{\tau} d\tau' \left( \begin{aligned} &[\mu_I(\tau), \rho_I(\tau') \mu_I(\tau')] G_X^+(x(\tau'), x(\tau)) \\ &- [\mu_I(\tau), \mu_I(\tau') \rho_I(\tau')] G_X^+(x(\tau), x(\tau')) \end{aligned} \right) + O(\mu^3)$$

- Return to the Schrödinger picture:

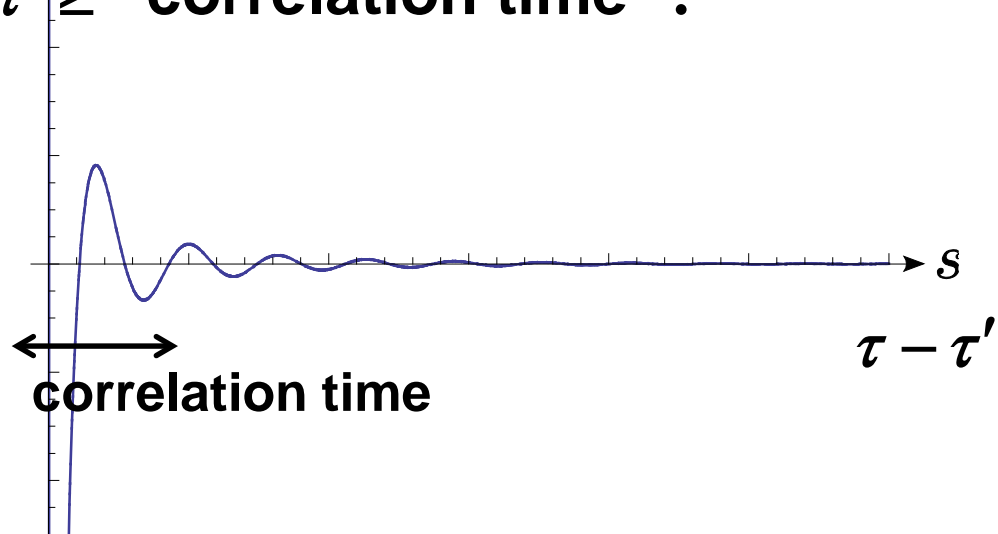
$$\begin{aligned} & \dot{\rho}_{mn}(\tau) + i(E_m - E_n) \rho_{mn}(\tau) \\ &= \sum_{k,\ell} \int_{\tau_1}^{\tau} d\tau' \left( \begin{array}{l} e^{-(E_m - E_k)(\tau' - \tau)} \mu_{mk} \mu_{\ell n} \rho_{k\ell}(\tau') G_X^+(x(\tau'), x(\tau)) \\ + e^{-(E_m - E_\ell)(\tau - \tau')} \mu_{mk} \mu_{\ell n} \rho_{k\ell}(\tau') G_X^+(x(\tau), x(\tau')) \\ - e^{-(E_\ell - E_m)(\tau' - \tau)} \mu_{k\ell} \mu_{\ell n} \rho_{mk}(\tau') G_X^+(x(\tau'), x(\tau)) \\ - e^{-(E_k - E_n)(\tau - \tau')} \mu_{mk} \mu_{k\ell} \rho_{\ell n}(\tau') G_X^+(x(\tau), x(\tau')) \end{array} \right) + O(\mu^3) \end{aligned}$$

The Wightman function  $G_X^+(x(\tau), x(\tau'))$  has a correlation in the time direction

and almost vanishes if  $\tau - \tau' \geq$  "correlation time".



$\rho_{k\ell}(t')$  in the integrand can be replaced by  $\rho_{k\ell}(t)$



Further assuming that off-diagonal elements can be neglected, we have a closed equation for the density distribution  $\rho_{mm}(t)$ :



(to be the case in examples below)

$$\dot{\rho}_{mm}(\tau) = \sum_{k \neq m} \left[ w_{mk}(\tau; \tau_1) \rho_{kk}(\tau) - w_{km}(\tau; \tau_1) \rho_{mm}(\tau) \right]$$

with  $w_{mk}(\tau; \tau_1) = |\mu_{mk}|^2 \dot{F}(E_m - E_k; \tau; \tau_1)$

$$\dot{F}(\Delta E; \tau; \tau_1) \equiv \int_{\tau_1}^{\tau} d\tau' \begin{bmatrix} e^{-i\Delta E(\tau' - \tau)} G_X^+(x(\tau'), x(\tau)) \\ + e^{-i\Delta E(\tau - \tau')} G_X^+(x(\tau), x(\tau')) \end{bmatrix}$$

**NB:** Suppose  $\dot{F}^{\text{eq}}(\Delta E) \equiv \lim_{\tau - \tau_1 \rightarrow \infty} \dot{F}(\Delta E; \tau; \tau_1)$  satisfies the relation

$$\dot{F}^{\text{eq}}(\Delta E) = e^{-\beta \Delta E} \dot{F}^{\text{eq}}(-\Delta E) \left( \Leftrightarrow \frac{w_{mk}(\tau; \tau_1)}{w_{km}(\tau; \tau_1)} = e^{-\beta(E_m - E_k)} \right)$$

$$\Rightarrow \text{Detailed balance: } \lim_{\tau - \tau_1 \rightarrow \infty} w_{mk}(\tau; \tau_1) \rho_{kk}^{\text{eq}} = \lim_{\tau - \tau_1 \rightarrow \infty} w_{km}(\tau; \tau_1) \rho_{mm}^{\text{eq}}$$

$$\Rightarrow \frac{\rho_{mm}^{\text{eq}}}{\rho_{kk}^{\text{eq}}} = e^{-\beta(E_m - E_k)} \Rightarrow \rho_{mm}^{\text{eq}} = \frac{1}{Z} e^{-\beta E_m}$$

## **6. Unruh-DeWitt detector in de Sitter space**

**[2] MF-Sakatani-Sugishita [PRD89 (2014) 064024, 1305.0256]**

## ■ Setup

- classical path

$$\left( ds^2 = \ell^2 \frac{-d\eta^2 + d\mathbf{x}^2}{\eta^2} \right)$$

$$x^\mu(\tau) = (\eta(\tau), \mathbf{x}(\tau)) = (-\ell e^{-\tau/\ell}, \mathbf{0})$$

- Wightman function

$$G^+(x, x'; \eta_0) = \langle \mathbf{0}_{\eta_0} | \phi_I(x) \phi_I(x') | \mathbf{0}_{\eta_0} \rangle$$

$$= \int \frac{d^{d-1}\mathbf{k}}{(2\pi)^{d-1}} G_k^+(\eta, \eta'; \eta_0) \cos \mathbf{k} \cdot (\mathbf{x} - \mathbf{x}')$$

- What we need to do:

(1) Calculate  $G_k^+(\eta, \eta'; \eta_0)$

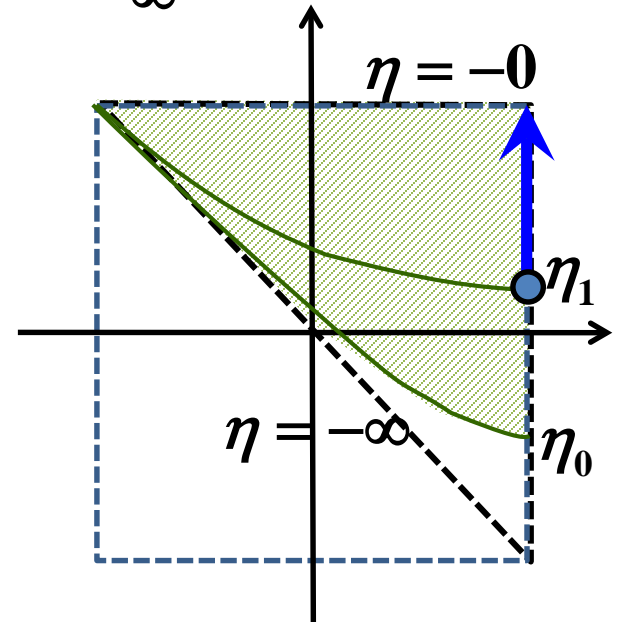
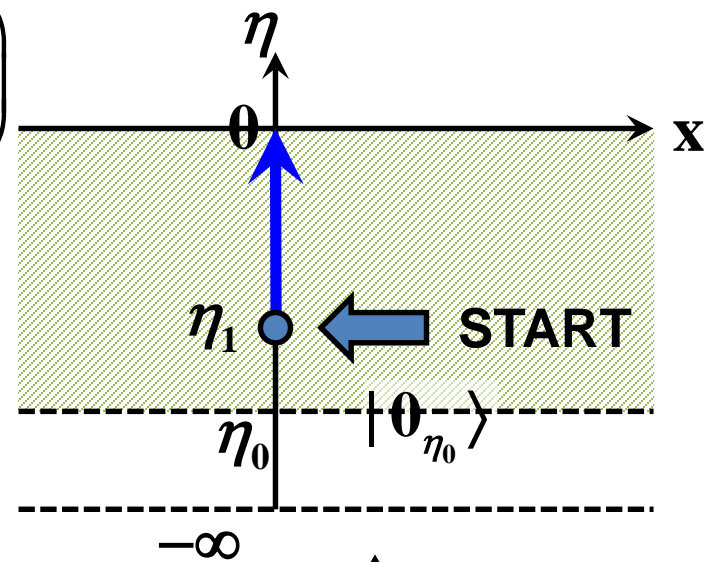
(2) Expand it around  $\eta_0 = -\infty$  ( $\Leftrightarrow$  BD)

(3) Integrate over  $k$  to obtain

$$\dot{F}(\Delta E; \tau; \tau_1; \eta_0 = -e^{-\tau_0})$$

$$\equiv \int_{\tau_1}^{\tau} d\tau' \left[ \begin{array}{l} e^{-i\Delta E(\tau'-\tau)} G_X^+(x(\tau'), x(\tau); \eta_0) \\ + e^{-i\Delta E(\tau-\tau')} G_X^+(x(\tau), x(\tau'); \eta_0) \end{array} \right] = \sum_{n \geq 0} e^{-n(\tau-\tau_0)} \dot{F}^{(n)}(\Delta E; \tau; \tau_1)$$

$(\tau_0 \equiv -\ln(-\eta_0))$



## ■ Result

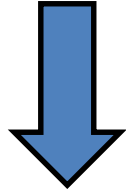
$$\begin{aligned} & \dot{F} \left( \Delta E ; \tau ; \tau_1 ; \eta_0 = -e^{-\tau_0} \right) \\ &= \dot{F}^{(0)} \left( \Delta E ; \tau ; \tau_1 \right) + e^{-2(\tau-\tau_0)} \dot{F}^{(2)} \left( \Delta E ; \tau ; \tau_1 \right) + \dots \end{aligned}$$

with

$$\begin{aligned} \dot{F}^{(0)} &= \frac{e^{-\pi \Delta E} \Gamma\left(\frac{d-1}{2} + \nu + i \Delta E\right) \Gamma\left(\frac{d-1}{2} - \nu + i \Delta E\right) \Gamma\left(\frac{d-1}{2} + \nu - i \Delta E\right) \Gamma\left(\frac{d-1}{2} - \nu - i \Delta E\right)}{8 \pi^{\frac{d+1}{2}} \Gamma\left(\frac{d-1}{2}\right)} \\ &+ \frac{e^{i \pi \frac{d-1}{2}} e^{i \pi \nu} e^{-(\frac{d-1}{2} + \nu - i \Delta E)(\tau - \tau_1)}}{8 \pi^{\frac{d-1}{2}} \Gamma\left(\frac{d-1}{2}\right) \sin(\pi \nu)} {}_3\widehat{F}_2\left(\begin{matrix} \frac{d-1}{2}, \frac{\frac{d-1}{2} + \nu - i \Delta E}{2}, \frac{d-1}{2} + \nu \\ 1 + \nu, \frac{\frac{d+3}{2} + \nu - i \Delta E}{2} \end{matrix}; e^{-2(\tau - \tau_1)}\right) \\ &+ \frac{e^{i \pi \frac{d-1}{2}} e^{-i \pi \nu} e^{-(\frac{d-1}{2} - \nu - i \Delta E)(\tau - \tau_1)}}{8 \pi^{\frac{d-1}{2}} \Gamma\left(\frac{d-1}{2}\right) \sin(-\pi \nu)} {}_3\widehat{F}_2\left(\begin{matrix} \frac{d-1}{2}, \frac{\frac{d-1}{2} - \nu - i \Delta E}{2}, \frac{d-1}{2} - \nu \\ 1 - \nu, \frac{\frac{d+3}{2} - \nu - i \Delta E}{2} \end{matrix}; e^{-2(\tau - \tau_1)}\right) \\ &+ \frac{e^{-i \pi \frac{d-1}{2}} e^{-i \pi \nu} e^{-(\frac{d-1}{2} + \nu + i \Delta E)(\tau - \tau_1)}}{8 \pi^{\frac{d-1}{2}} \Gamma\left(\frac{d-1}{2}\right) \sin(\pi \nu)} {}_3\widehat{F}_2\left(\begin{matrix} \frac{d-1}{2}, \frac{\frac{d-1}{2} + \nu + i \Delta E}{2}, \frac{d-1}{2} + \nu \\ 1 + \nu, \frac{\frac{d+3}{2} + \nu + i \Delta E}{2} \end{matrix}; e^{-2(\tau - \tau_1)}\right) \\ &+ \frac{e^{-i \pi \frac{d-1}{2}} e^{i \pi \nu} e^{-(\frac{d-1}{2} - \nu + i \Delta E)(\tau - \tau_1)}}{8 \pi^{\frac{d-1}{2}} \Gamma\left(\frac{d-1}{2}\right) \sin(-\pi \nu)} {}_3\widehat{F}_2\left(\begin{matrix} \frac{d-1}{2}, \frac{\frac{d-1}{2} - \nu + i \Delta E}{2}, \frac{d-1}{2} - \nu \\ 1 - \nu, \frac{\frac{d+3}{2} - \nu + i \Delta E}{2} \end{matrix}; e^{-2(\tau - \tau_1)}\right) \end{aligned}$$

$$\dot{F}^{(2)} \left( \Delta E ; \tau ; \tau_1 \right) = \dots$$

We are interested in time scales larger than  $1 / \Delta E$ .



Taking an average over  $\tau_1$   
with the duration  $\Delta\tau \simeq 1 / \Delta E$

$$\begin{aligned} \dot{F}(\Delta E; \tau; \eta_0 = -e^{-\tau_0}) &\equiv \frac{1}{\Delta\tau} \int_{\tau_1 - \Delta\tau/2}^{\tau_1 + \Delta\tau/2} d\tau'_1 \dot{F}(\Delta E; \tau; \tau'_1; \eta_0 = -e^{-\tau_0}) \\ &= \dot{F}^{\text{eq}}(\Delta E) + e^{-2(\tau - \tau_0)} \dot{F}^{(2)}(\Delta E) + O(e^{-4(\tau - \tau_0)}) \end{aligned}$$

with

$$\begin{aligned} \dot{F}^{\text{eq}}(\Delta E) &= \frac{e^{-\pi\ell\Delta E}}{8\pi^{(d+1)/2} \Gamma\left(\frac{d-1}{2}\right)} \Gamma\left(\frac{d-1+2(\nu+i\ell\Delta E)}{4}\right) \Gamma\left(\frac{d-1+2(-\nu+i\ell\Delta E)}{4}\right) \\ &\quad \times \Gamma\left(\frac{d-1+2(\nu-i\ell\Delta E)}{4}\right) \Gamma\left(\frac{d-1-2(\nu+i\ell\Delta E)}{4}\right) \end{aligned}$$

$$\begin{aligned}
\dot{\mathcal{F}}^{(2)}(\Delta E) &= \left(\frac{d-2}{4}\right)^2 \frac{1}{2^{3+n} \pi^{\frac{d-1}{2}} \Gamma\left(\frac{d-1}{2}\right)} \\
&\times \left[ iG_{3,3}^{2,2} \left( e^{i\pi} \left| \begin{array}{c} -\frac{d-3-n}{2}, -\frac{\frac{d-5}{2}-\nu-n-i\Delta E}{2}, -\frac{d-3-n}{2} + \nu \\ 0, \nu, -\frac{\frac{d-1}{2}-\nu-n-i\Delta E}{2} \end{array} \right. \right) \right. \\
&\quad - iG_{3,3}^{2,2} \left( e^{-i\pi} \left| \begin{array}{c} -\frac{d-3-n}{2}, -\frac{\frac{d-5}{2}-\nu-n-i\Delta E}{2}, -\frac{d-3-n}{2} + \nu \\ 0, \nu, -\frac{\frac{d-1}{2}-\nu-n-i\Delta E}{2} \end{array} \right. \right) \\
&\quad - G_{4,4}^{2,3} \left( e^{i\pi} \left| \begin{array}{c} -\frac{d-3-n}{2}, -\frac{d-3-n}{2} + \nu, -\frac{\frac{d-5}{2}-\nu-n-i\Delta E}{2}, -\frac{d-4-n}{2} + \nu \\ 0, \nu, -\frac{d-4-n}{2} + \nu, -\frac{\frac{d-1}{2}-\nu-n-i\Delta E}{2} \end{array} \right. \right) \\
&\quad \left. - G_{4,4}^{2,3} \left( e^{-i\pi} \left| \begin{array}{c} -\frac{d-3-n}{2}, -\frac{d-3-n}{2} + \nu, -\frac{\frac{d-5}{2}-\nu-n-i\Delta E}{2}, -\frac{d-4-n}{2} + \nu \\ 0, \nu, -\frac{d-4-n}{2} + \nu, -\frac{\frac{d-1}{2}-\nu-n-i\Delta E}{2} \end{array} \right. \right) \right] \\
&\quad + (\Delta E \rightarrow -\Delta E).
\end{aligned}$$

$$\left( G_{p,q}^{m,n} \left( z \left| \begin{array}{c} a_1, \dots, a_p \\ b_1, \dots, b_q \end{array} \right. \right) : \text{Meijer's } G\text{-function} \right)$$



## ■ Properties

$$(1) \dot{F}^{\text{eq}}(\Delta E) = e^{-2\pi\ell\Delta E} \dot{F}^{\text{eq}}(-\Delta E)$$

$$\Rightarrow \rho_{mm}^{\text{eq}} = \frac{1}{Z} e^{-2\pi\ell E_m} \left( \text{i.e. } T = \frac{1}{\beta} = \frac{1}{2\pi\ell} \right)$$

$$(2) A(\Delta E) \equiv \dot{F}^{\text{eq}}(\Delta E) + \dot{F}^{\text{eq}}(-\Delta E)$$

$$\simeq \frac{|\Delta E|^{d-3}}{2^{d-2} \pi^{(d-3)/2} \Gamma((d-1)/2)} \quad (|\Delta E| \ell \rightarrow +\infty)$$

$$(3) \dot{F}^{(2)}(\Delta E) = \dot{F}^{(2)}(-\Delta E)$$

## ■ Example: two-level detector

Let us consider a two-level system with

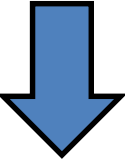
(1)  $\Delta E \equiv E_2 - E_1 > 0$

(2)  $\Delta E \gg 1 / \ell$

(3)  $\mu = (\mu_{mn}) = \begin{pmatrix} 0 & \mu_{12} \\ \mu_{12}^* & 0 \end{pmatrix}$

 master equation is decomposed  
to the diagonal and off-diagonal parts

(4) detector is "ideal" :  $\lambda^2 |\mu_{12}|^2 (\Delta E)^{d-3} \ell \gg 1$

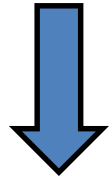


$$\begin{pmatrix} \dot{\rho}_{11}(\tau) \\ \dot{\rho}_{22}(\tau) \end{pmatrix} = \begin{pmatrix} -w_+(\tau; \tau_0) & w_-(\tau; \tau_0) \\ w_+(\tau; \tau_0) & -w_-(\tau; \tau_0) \end{pmatrix} \begin{pmatrix} \rho_{11}(\tau) \\ \rho_{22}(\tau) \end{pmatrix}$$

with  $w_{\pm}(\tau; \tau_0) \equiv \lambda^2 |\mu_{12}|^2 \left[ \dot{F}^{\text{eq}}(\pm \Delta E) + e^{-2(\tau - \tau_0)/\ell} \dot{F}^{(2)}(\pm \Delta E) \right]$

Expand  $\rho_{mm}(\tau)$  around the equilibrium distribution:

$$\rho_{mm}(\tau) = \rho_{mm}^{\text{eq}} + \Delta\rho_{mm}(\tau) \left( \rho_{mm}^{\text{eq}} = \frac{e^{-e\pi\ell E_{mm}}}{e^{-e\pi\ell E_{11}} + e^{-e\pi\ell E_{22}}}; m = 1, 2 \right)$$



$$\Delta\rho_{11}(\tau) + \Delta\rho_{22}(\tau) = 0$$

$$\Delta\rho_{11}(\tau)$$

$$= e^{-\lambda^2 |\mu_{12}|^2 A(\Delta E)(\tau - \bar{\tau}_1)} \Delta\rho_{11}(\bar{\tau}_1) \quad (\bar{\tau}_1 \approx \tau_1)$$

$$- \frac{\lambda^2 |\mu_{12}|^2 \dot{F}^{(2)}(\Delta E)}{\lambda^2 |\mu_{12}|^2 A(\Delta E) - 2} \tanh(\pi\ell\Delta E) \left( e^{-2(\tau - \bar{\tau}_1)/\ell} - e^{-\lambda^2 |\mu_{12}|^2 A(\Delta E)(\tau - \bar{\tau}_1) - 2(\bar{\tau}_1 - \tau_0)/\ell} \right)$$

$$\left( \text{NB: } A(\Delta E) = \dot{F}^{\text{eq}}(\Delta E) + \dot{F}^{\text{eq}}(-\Delta E) \sim |\Delta E|^{d-3} \right)$$

$\therefore$  For an ideal detector,

$$\Delta\rho_{11}(\tau) = - \frac{\dot{F}^{(2)}(\Delta E)}{A(\Delta E)} \tanh(\pi\ell\Delta E) e^{-2(\tau - \bar{\tau}_1)/\ell}$$

## ■ Properties

$$\Delta\rho_{11}(\tau) = -\frac{\dot{F}^{(2)}(\Delta E)}{A(\Delta E)} \tanh(\pi\ell\Delta E) e^{-2(\tau-\bar{\tau}_1)/\ell}$$

- Everything is in a universal form.

(i.e., the expression does not depend  
on details of the detector.)

- The relaxation time is given by the universal quantity,

$$\tau_{\text{relax}} = \frac{\ell}{2}$$

This gives an example of  
the nonequilibrium-thermodynamic quantities  
intrinsic to de Sitter space.

## **7. Conclusion and outlook**

## ■ What we have done:


### [PART 1]

- Proposed to take the vacuum in curved spacetime to be the set of instantaneous ground states.
- Constructed and developed a framework to calculate various propagators for the vacuum.

### [PART 2]

- Wrote down the master equation which completely determines a finite time evolution of the density matrix of an Unruh-DeWitt detector in  $\forall$  spacetime.
- Applied the framework to uncover a universal nonequilibrium property intrinsic to de Sitter space.

## ■ Future direction

- (1) It should be important to establish the framework for interacting quantum fields, especially for those in de Sitter space.**
- (2) ----- to apply the present framework to investigate the cosmological constant problem.**
- (3) ----- to extend the framework to cases where a horizon exists.**
  -  This may be helpful in understanding the thermodynamic aspects of dynamical BH.**
- (4) ...**

**Thank you.**