Fast scrambling in holographic Einstein-Podolsky-Rosen pair

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> > 特任助教 from Jan.

"Fast scrambling in holographic Einstein-Podolsky-Rosen pair", arXiv:1708.09493 [hep-th], JHEP 1711 (2017) 049



My research interest

Classical and quantum aspects of black holes, and their applications. Fast scrambling in holographic Einstein-Podolsky-Rosen pair

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1. Introduction

Fast scrambling conjecture

"Black holes are the fastest scramblers in nature" Sekino&Susskind, 08



The electric charge spreads out all over the horizon.

Its time scale:



Scrambling = Delocalization of local quantum excitation

What is scrambling time of a conventional matter? diffusion

coefficient

d-dim

Scrambling time:

Number of $\mathsf{DOF}\infty$ Volume

Scale of the excitation: $\langle \Delta x^2
angle \sim \kappa t$ $t_* \sim \frac{1}{\kappa} L^2$ $L^d \propto N$

• $t_* \propto N^{2/d}$

BH's scrambling is much faster than conventional matters. $\ln N \ll N^{2/d}$

Quantum DOF of BH would be strongly chaotic.

Fast scrambling in AdS/CFT

Shenker&Stanford, 13



Thermal property is explained by the entanglement.

$$\langle \Psi | \mathcal{O}_R | \Psi \rangle = Z^{-1} \mathrm{tr}[e^{-\beta H_R} \mathcal{O}_R]$$

Correlation is a measure of scrambling. Shenker&Stanford, 13

Left CFT

Right CFT

Assume that CFT is "chaotic".

but entangled

causally disconnected

Give a perturbation on Left CFT.

Left CFT is scrambled by "chaos".



Destroy the correlation between L and R

The correlation between left and right CFTs quantifies the scrambling.

present section, we will consider a more physical system, by numerica ection of thermal qubits. Although we are limited to a rather small system will be visible. Shenker&Stanford, 13 Jsing sparse marix techques Dieposible to Sinste Velpure state by qubits. We will be less ambitious, studying a system (L) made up 10 spins another ten for the thermofield double (R), We will use an Ising Han n transverse and parallel magnetic fields: Hamiltonian: $H_{L} = \sum_{\substack{i=1 \\ H_{L}^{i} = 1}} \left\{ \sigma_{z}^{(i)} \sigma_{z}^{(i+1)} - 1.05 \ \sigma_{x}^{(i)} + 0.5 \ \sigma_{z}^{(i)} \right\}.$ $H_{L}^{i=1} = \sum_{\substack{i=1 \\ T_{L}^{i} = 1}} \left\{ \sigma_{z}^{(i)} \sigma_{z}^{(i+1)} - 1.05 \ \sigma_{x}^{(i)} + 0.5 \ \sigma_{z}^{(i)} \right\}.$ $|\Psi\rangle = \frac{1}{\sqrt{Z}} \sum e^{-\beta E_n/2} |n\rangle_L |n\rangle_R .$ Thermofield double: coefficients -1.05 and 0.5 are chosen, following [31], to ensure that the Give a roment satisfication $|\Psi'\rangle = \sigma_z^{(5,L)}(-t_w)|\Psi\rangle$ Decay of the Durat procedure is to prepare the thermofield double state at the in $\operatorname{correlation} \operatorname{at}^{\mathsf{be}} = 0$. We then apply a perturbation $\sigma_{\lambda}^{(5,L)}$ to the fifth $\sigma_{\lambda}^{(5,L)}$ em ab: a time t_w in the past. In other words, we consider the perturbed



Bound on "Lyapunov" exponent

Maldacena, Shencker&Stanford, 16

Correlation function ~ OTOC



perturbed thermofield double



Lyapunov bound Maldacena, Shenker&Stanford, 16

- 1. Quantum Lyapunov exponent is defined by the OTOC. $\langle W(t)V(0)W(t)V(0)\rangle_{\beta} \sim 1 - \# e^{\lambda_L t}$
- 2. The Lyapunov exponent has upper bound. $\lambda_L \leq 2\pi T$ AdS/CFT $\longrightarrow \lambda_L = 2\pi T$ "refinement of the fast scrambling conjecture"

3. If a QFT saturates the bound, it has an Einstein dual. most chaotic black hole dual

(ex. SYK model)

In this talk

We propose one of the simplest model which shows fast scrambling.

Holographic EPR pair.

quark-antiquark in N=4 SYM



entangled causally disconnected

$$\lambda_L = 2\pi T \quad \tau_* \sim \beta \ln S$$

Dual picture of the EPR pair is just a probe string.



Fast scrambling or Lyapunov bound do not directly imply the existence of an Einstein dual.

Holographic EPR pair

Setup

We consider a probe classical string in AdS5.

$$ds^{2} = \frac{\ell^{2}}{z^{2}} \left(-dt^{2} + dz^{2} + dx^{2} + dy_{1}^{2} + dy_{2}^{2}\right) ,$$

$$S_{\rm NG} = -\frac{1}{2\pi\alpha'} \int d^{2}\sigma \sqrt{-h} ,$$

For simplicity, we consider the string in y1=y2=0 plane. (Effectively AdS3)

Mikhailov, 03

Mikhailov solution

Worldsheet coordinates: (au, σ)

$$\begin{split} t &= \frac{\dot{f}_t(\tau)}{\sigma} + f_t(\tau) \ , \quad x = \frac{\dot{f}_x(\tau)}{\sigma} + f_x(\tau) \ , \quad z = \frac{1}{\sigma} \ , \\ \text{where} \quad -\dot{f}_t^2 + \dot{f}_x^2 = -1 \ . \end{split}$$

$$t = f_t(\tau), \ x = f_x(\tau)$$

X

AdS boundary: $\sigma = \infty$.

au : proper time of the end point

Non-linear wave from AdS boundary

$$t = \frac{\dot{f}_t(\tau)}{\sigma} + f_t(\tau), \quad x = \frac{\dot{f}_x(\tau)}{\sigma} + f_x(\tau), \quad z = \frac{1}{\sigma}.$$
Holographic EPR solution

As a special case
$$f_t(\tau) = \frac{1}{a} \sinh a\tau, \quad f_x(\tau) = -\frac{1}{a} \cosh a\tau.$$

$$x^2 + z^2 = t^2 + \frac{1}{a^2}$$
string profile
$$\vec{q} = \vec{q}$$
String profile

Causal structure of the worldsheet

Induced metric $ds_h^2 = -[\sigma^2 - a^2]d\tau^2 + 2d\tau d\sigma \ .$ $U = e^{a\tau} \ , \quad V = -\frac{\sigma - a}{\sigma + a}e^{-a\tau}$ $ds_h^2 = -\frac{4dUdV}{(1 + UV)^2} \ .$

$$t = \frac{1}{a} \frac{U+V}{1-UV}$$
, $x = -\frac{1}{a} \frac{U-V}{1-UV}$, $z = \frac{1}{a} \frac{1+UV}{1-UV}$

Right quark "appears" by the maximal extension.

Why is this solution called EPR pair?



"Causally disconnected" & "entanglement" = "EPR pair"

Unruh temperature and Entropy

Accelerating particle feels thermal bath. Unruh effect

$$T = \beta^{-1} = \frac{a}{2\pi} \qquad \text{Unruh temperature}$$

Accelerating quark has entropy & energy.

Jensen&Karch, 14 acceleration: a_{μ}

$$E = \frac{4}{3}\sqrt{\lambda}T .$$

$$\lambda: \text{'t hooft coupling}$$

$$F = -TS_{\text{on-shell}}$$

$$S = dF/dT$$

$$E = F + TS$$

$$A: \text{'t hooft coupling}$$

$$Cloud of gluons$$

$$Cloud of gluon$$

Holographic EPR with a shock

$$t = \frac{\dot{f}_t(\tau)}{\sigma} + f_t(\tau) , \quad x = \frac{\dot{f}_x(\tau)}{\sigma} + f_x(\tau) , \quad z = \frac{1}{\sigma} .$$

Changing the acceleration

As a perturbation, we slightly change the acceleration at $\tau=\tau_0$.



junction condition:

$$c_1 = -(a'-a)\tau_0$$
, $c_2 = \left(\frac{1}{a} - \frac{1}{a'}\right)\sinh a\tau_0$, $c'_2 = \left(\frac{1}{a} - \frac{1}{a'}\right)\cosh a\tau_0$.



String dynamics

Even for tiny change of the acceleration, the string profile is significantly deformed.

induced metric of the string $ds_h^2 = -[\sigma^2 - M(\tau)]d\tau^2 + 2d\tau d\sigma ,$ $M(\tau) = a^2 + (a'^2 - a^2)\theta(\tau - \tau_0) .$

Vaidya-like geometry with infalling null shock.



Fast scrambling of EPR

Correlation between quark and antiquark



geodesic distance

$$\left(\delta a \to 0 , \quad \tau_0 \to -\infty , \quad \gamma \equiv \frac{\delta a}{2a} e^{-a\tau_0} : \text{fixed } . \right)$$

Visualizing the correlation



earlier change of acceleration

Scrambling time

$$\langle F_L(0)F_R(0)\rangle_W \sim \left(1 + \frac{\delta a}{4a}e^{a|\tau_0|}\right)^{-2}$$



Scrambling time:

$$\tau_* \sim \frac{1}{a} \ln\left(\frac{a}{\delta a}\right)$$

2

•



Holographic EPR is a fast scrambler.

Saturate Lyapunov bound



$$\cdot F_{L}F_{R}\cdot_{W} \sim \left(1+\frac{\delta a}{4a}e^{a|\tau_{0}|}\right)^{-2}$$

saturates the bound.

Saturation of the Lyapunov bound Einstein dual.

Summary

We proposed one of the simplest model showing fast scrambling: holographic EPR pair.

scrambling time

Lyapunov exponent $\lambda_L = 2\pi T$

Fast scrambler Saturation of the Lyapunov bound



(SYK is really Einstein gravity?)





Fast scrambling if tau1-tau2 is not too small.

Decreasing the accerelation

$$\delta a < 0 \qquad \longrightarrow \qquad \gamma < 0$$
$$\langle F_L F_R \rangle_W \sim \left(1 + \frac{\gamma}{2} \right)^{-2} = \left(1 - \frac{|\delta a|}{4a} e^{a|\tau_0|} \right)$$

Diverging? What is happening?

One-way traverse wormhole

Causally connected. "One-way traversable wormhole"

Divergence of the correlation or OTOC.

induced metric

$$ds_h^2 = -[\sigma^2 - M(\tau)]d\tau^2 + 2d\tau d\sigma ,$$

$$M(\tau) = a^2 + (a'^2 - a^2)\theta(\tau - \tau_0) .$$

Vaidya-like geometry with infalling null shock.

Shock surface

Note $\delta a = a' - a$ can be negative.

On the other hand, for Vaidya spacetime, the deviation of the mass must be positive.

Correlation between quark and antiquark

Correlation: $\langle F_L(\tau_L)F_R(\tau_R)\rangle_W \sim e^{-d}$

Field on worldsheet is the target space coords. Its conjugate operator is the force. $x(\tau, \sigma)$

$$\mathcal{L}_{\text{quark}} \sim -m\sqrt{1-v^2} + V(x)$$
$$\frac{\partial \mathcal{L}_{\text{quark}}}{\partial x} \sim F$$

geodesic

distance

geodesic distance

$$\left\{ \delta a \to 0 , \quad \tau_0 \to -\infty , \quad \gamma \equiv \frac{\delta a}{2a} e^{-a\tau_0} : \text{fixed} . \right\}$$

Comment on the singularity

The induced metric is regular here.

 $ds_h^2 = -\frac{4dUdV}{(1+UV)^2} \; .$

But the string solution diveges

$$t = \frac{1}{a} \frac{U+V}{1-UV}$$
, $x = -\frac{1}{a} \frac{U-V}{1-UV}$, $z = \frac{1}{a} \frac{1+UV}{1-UV}$.

The "singularity" corresponds to the Poincare horizon in the target space.

If we take the global coordinates for the target space, there is no divergence.

Spacetime structure

For
$$\mathcal{T} < \mathcal{T}_{0}$$

we introduced (U,V)-coordinates.
 $U = e^{a\tau}$, $V = -\frac{\sigma - a}{\sigma + a}e^{-a\tau} \iff \tau = \frac{1}{a}\ln U$, $\sigma = a\frac{1 - UV}{1 + UV}$. $U = e^{a\tau_{0}}$
For $\mathcal{T} > \mathcal{T}_{0}$
we introduce (U',V')-coordinates as
 $U' = e^{a'\tau}$, $V' = -\frac{\sigma - a'}{\sigma + a'}e^{-a'\tau}$
 $\iff \tau = \frac{1}{a'}\ln U'$, $\sigma = a'\frac{1 - U'V'}{1 + U'V'}$.
 $U' = e^{a'\tau_{0}}$

glue

Matching condition at shock surface?

Matching condition $U = e^{a\tau} , \quad V = -\frac{\sigma - a}{\sigma + a} e^{-a\tau} \iff \tau = \frac{1}{a} \ln U , \quad \sigma = a \frac{1 - UV}{1 + UV} .$ $U' = e^{a'\tau} , \quad V' = -\frac{\sigma - a'}{\sigma + a'} e^{-a'\tau} \iff \tau = \frac{1}{a'} \ln U' , \quad \sigma = a' \frac{1 - U'V'}{1 + U'V'} .$ (τ, σ) is common in (U,V) and (U',V') patches. $a \frac{1 - UV}{1 + UV}\Big|_{\tau = \tau_0} = a' \frac{1 - U'V'}{1 + U'V'}\Big|$ $\frac{a+a') e^{a\tau_0} V}{\delta a \, e^{a\tau_0} \, V + (a+a')} e^{-a'\tau_0} \, .$

$$\delta a \to 0 , \qquad \tau_0 \to -\infty , \qquad \gamma \equiv \frac{\delta a}{2a} e^{-a\tau_0} : \text{ fixed } .$$

$$\frac{a + a') e^{a\tau_0} V}{\delta a e^{a\tau_0} V + (a + a')} e^{-a'\tau_0} . \qquad \checkmark' = V + \gamma ,$$

 $U = e^{a\tau_0} \qquad \qquad U' = e^{a'\tau_0}$

OTOC measures chaos?

Maldacena, Shenker & Stanford, 16

$$C_T = -\langle [\hat{x}(t), \hat{p}(0)]^2 \rangle \sim \{x(t), p(0)\}_{\rm PB}^2$$

$$= \left(\frac{\delta x(t)}{\delta x(0)}\right)^2 \quad \to e^{2\lambda t}$$

 λ :Lyapunov

What is scrambling

Scrambling = delocalization of local quantum information

Scrambling ~ Thermalization ~ Quantum chaos, Butterfly effect

Important in the context of black hole physics or AdS/CFT.

Validity of the geodesic approximation

 $\langle F_L F_R \rangle_W \sim e^{-\Delta d}$

is effective for large Δ .

Our case is $\Delta = 1...$

We use the geodesic approximation just for the rough estimation of the correlation.

(For pure AdS, the geodesic approximation is exact for any Δ .)

What is scrambling

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Fast scrambling conjecture

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In Schwarzshild coordinates, the field line surrounds the BH.

We cannot distinguish BH and charge

 $t_* \sim \beta \ln S \sim \beta \ln N$

What is scrambling time of a conventional matter? diffusion

coefficient

d-dim

Scrambling time:

Number of $\mathsf{DOF}\infty$ Volume

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• $t_* \propto N^{2/d}$

BH's scrambling is much faster than conventional matters. $\ln N \ll N^{2/d}$

Quantum DOF of BH would be strongly chaotic. $t_* \sim \beta \ln S$: Fast scrambler

Lyapunov bound Maldacena, Shencker&Stanford, 16 1. Quantum Lyapunov exponent is defined by the OTOC. $\langle W(t)V(0)W(t)V(0)\rangle_{\beta} \sim 1 - \#e^{\lambda_L t}$ $\int C_T = -\langle [\hat{x}(t), \hat{p}(0)]^2 \rangle \sim \{x(t), p(0)\}_{PB}^2 = \left(\frac{\delta x(t)}{\delta x(0)}\right)^2 \rightarrow e^{2\lambda t}$

2. The Lyapunov exponent has upper bound. $\lambda_L \leq 2\pi T$ AdS/CFT $\implies \lambda_L = 2\pi T$

"refinement of the fast scrambling conjecture"

3. If a QFT saturates the bound, it has an Einstein dual.

(ex. SYK model)

OTOC

saturates the bound.

Saturation of the Lyapunov bound Einstein dual.