

The muon g-2: a new data-based analysis

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Partially based on

A. Keshavarzi, DN and T. Teubner **(KNT)**

arXiv:1802.02995 (Phys. Rev. D97 (2018) 114025) **(KNT18)**

arXiv:1911.00367 (Phys. Rev. D101 (2020) 014029) **(KNT19)**

Muon g-2: introduction

Lepton magnetic moment $\vec{\mu}$: $\mathcal{H} = -\vec{\mu} \cdot \vec{B}$

$$\boxed{\vec{\mu} = -g \frac{e}{2m} \vec{s}}, \quad (\vec{s} = \frac{1}{2} \vec{\sigma} \text{ (spin)}, \quad g = 2 + 2F_2(0))$$

where

$$\bar{u}(p+q)\Gamma^\mu u(p) = \bar{u}(p+q) \left(\gamma^\mu F_1(q^2) + \frac{i\sigma^{\mu\nu} q_\nu}{2m} F_2(q^2) \right) u(p)$$

Anomalous magnetic moment: $a \equiv (g - 2)/2$ ($= F_2(0)$)

Historically,

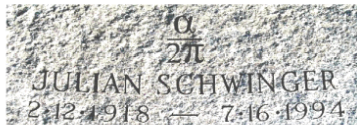
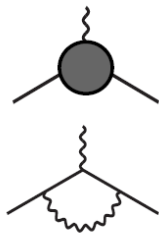
- ★ $g = 2$ (tree level, Dirac)
- ★ $a = \alpha/(2\pi)$ (1-loop QED, Schwinger)

Today, still important, since...

- ★ One of the **most precisely measured** quantities:

$$\boxed{a_\mu^{\text{exp}} = 11\,659\,208.9(6.3) \times 10^{-10}} \quad [0.5\text{ppm}] \quad (\text{Bennett et al})$$

- ★ **Extremely useful** in probing/constraining physics beyond the SM



Why Muon $g-2$?

- $\gtrsim 3.5 \sigma$ Anomaly Observed
Long standing anomaly (~ 20 yrs), in spite of careful studies on every aspect.
(\rightarrow Major theoretical blunder unlikely.)
Hint of New Physics beyond the Standard Model?
- No new physics at the LHC so far
Intensity frontier: more and more important
- Long history of research
1st $(g - 2)_\mu$ exp.: Garwin, Lederman & Weinrich (1957)
Well-established place to search for new physics
- Leptonic observable
Experimentally and theoretically clean

Muon g-2: previous exp. (after 1960)

Experiment	Years	Polarity	$a_\mu \times 10^{10}$	Precision [ppm]	Sensitivity
CERN I	1961	μ^+	11 450 000(220 000)	4300	2-loop QED contrib. (3600 ppm)
CERN II	1962-1968	μ^+	11 661 600(3100)	270	3-loop QED contrib. (260 ppm)
CERN III	1974-1976	μ^+	11 659 100(110)	10	hadronic vacuum polarization contrib. (60 ppm)
CERN III	1975-1976	μ^-	11 659 360(120)	10	
BNL	1997	μ^+	11 659 251(150)	13	
BNL	1998	μ^+	11 659 191(59)	5	4-loop QED contrib. (3.3 ppm)
BNL	1999	μ^+	11 659 202(15)	1.3	electroweak contrib. (1.3 ppm)
BNL	2000	μ^+	11 659 204(9)	0.73	hadronic light-by-light contrib. (0.86 ppm)
BNL	2001	μ^-	11 659 214(9)	0.72	hadronic NLO vacuum pol. contrib. (-0.85 ppm)
Average			11 659 208.0(6.3)	0.54	

Table from BNL-E821 final report, Phys. Rev. D 73 (2006) 072003

History of muon g-2 exp. is a history of SM tests.

This is not the whole story: the history still goes on.

Muon g-2 vs New Physics

Basically, **any** new particle which couples to the muon gives a non-zero contribution to the muon g-2:

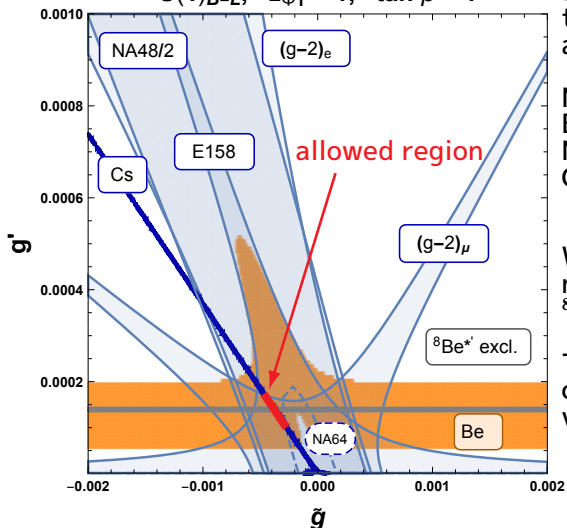
- SUSY particles ($\tilde{\mu}, \tilde{W}^{\pm}, \tilde{Z}^0, \tilde{B}^0, \dots$)
- extra Higgses ($H^{\pm}, A^0, H^{\pm\pm}, \dots$)
- Kaluza-Klein excitations of μ and γ
- extra Z -like particle ($Z', \text{"dark } Z", \dots$)
- extra γ -/axion- like light particle ("dark photon", ...)
- leptoquarks
- \vdots

In many cases, the mass and couplings of these new particles are free parameters. By tuning them, one can explain the muon g-2 anomaly. But it is often non-trivial to explain why Nature chooses such a parameter set.

E.g., Family universal type-I 2HDM: Allowed region

Allowed region in the (\tilde{g}, g') plane:

$U(1)_{B-L}, z_{\Phi 1} = 1, \tan \beta = 1$



Just an example when the $U(1)'$ charges = B-L and $\tan \beta = 1$

NA48/2: $\pi^0 \rightarrow Z'\gamma$ searches
 E158: Møller scattering
 NA64: e^- beam dump exp.
 Cs: atomic parity violation in Cs atom

White region is excluded by non-observation of Z' in ${}^8\text{Be}^{*'} \rightarrow {}^8\text{Be}$ transition

The strongest constraint comes from atomic parity violation in Cs.

Viable parameter region still exists.

Fig. from L. Delle Rose et al, arXiv:1812.05497

Breakdown of SM prediction for muon g-2

	<u>2011</u>	→	<u>2018</u>	<u>2019</u>
QED	11658471.81 (0.02)	→	11658471.90 (0.01) [arXiv:1712.06060]	
EW	15.40 (0.20)	→	15.36 (0.10) [Phys. Rev. D 88 (2013) 053005]	
LO HLbL	10.50 (2.60)	→	9.80 (2.60) [EPJ Web Conf. 118 (2016) 01005]	9.34 (2.92)
NLO HLbL			0.30 (0.20) [Phys. Lett. B 735 (2014) 90]	
	<u>HLMNT11</u>		<u>KNT18</u>	<u>KNT19</u>
LO HVP	694.91 (4.27)	→	693.27 (2.46) this work	692.78 (2.42)
NLO HVP	-9.84 (0.07)	→	-9.82 (0.04) this work	-9.83 (0.04)
NNLO HVP			1.24 (0.01) [Phys. Lett. B 734 (2014) 144]	
Theory total	11659182.80 (4.94)	→	11659182.05 (3.56) this work...	181.08 (3.78)
Experiment			11659209.10 (6.33) world avg	
Exp - Theory	26.1 (8.0)	→	27.1 (7.3) this work	28.0 (7.4)
Δa_μ	3.3 σ	→	3.7 σ this work	3.8 σ

(HVP: Hadronic Vacuum Polarization)
(HLbL: Hadronic Light-by-Light)

(Numbers taken from KNT18
and from KNT19)

QED contribution

QED contribution:

$$\begin{aligned} a_\mu(\text{QED}) &= \frac{\alpha}{2\pi} + 0.765857425(17) \left(\frac{\alpha}{\pi}\right)^2 + 24.05050996(32) \left(\frac{\alpha}{\pi}\right)^3 \\ &\quad + 130.8796(63) \left(\frac{\alpha}{\pi}\right)^4 + 753.3(1.0) \left(\frac{\alpha}{\pi}\right)^5 + \dots \\ &= 11658471.895(0.008) \times 10^{-10}, \quad (\text{numbers from PDG 2018}) \end{aligned}$$

where the uncertainty is dominated by that of α .

- 5-loop calculation! (Aoyama, Hayakawa, Kinoshita & Nio)
- The 4-loop corrections $\simeq 38 \times 10^{-10} \simeq \mathcal{O}(a_\mu(\text{exp}) - a_\mu(\text{SM}))$.
- The 4-loop contribution now fully cross-checked by another group. Mass-independent part by S. Laporta (*Phys.Lett.* **B772** (2017) 232), and mass-dependent part by A. Kurz et al (*Nucl. Phys.* **B879** (2014) 1; *Phys. Rev.* **D92** (2015) 073019; *ibid.* **D93** (2016) 053017)
- The 5-loop contribution very small ($\simeq 0.5 \times 10^{-10} \ll a_\mu(\text{exp}) - a_\mu(\text{SM})$)

Electroweak Contribution

Electroweak (EW) contribution:

$$a_{\mu}(\text{EW}) = \underbrace{19.48 \times 10^{-10}}_{\text{1-loop}} + \underbrace{(-4.12(10) \times 10^{-10})}_{\text{2-loop}} + \underbrace{\mathcal{O}(10^{-12})}_{\text{leading log 3-loop}}$$
$$= 15.36(10) \times 10^{-10}, \quad (\text{Number taken from PDG 2018})$$

where the uncertainty mainly comes from quark loops.

- 1-loop result published by many groups (Bardeen-Gastmans-Lautrup, Altarelli-Cabibbo-Maiani, Jackiw-Weinberg, Bars-Yoshimura, Fujikawa-Lee-Sanda) in 1972, and now a textbook exercise (Peskin & Schroeder's textbook, Problems 6.3 (Higgs) and 21.1 (W, Z))
- 2-loop contribution (~ 1700 diagrams in the 't Hooft-Feynman gauge) enhanced by $\ln(m_Z/m_{\mu})$ and also by a factor of $\mathcal{O}(10)$,

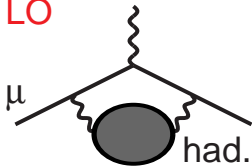
$$a_{\mu}(\text{EW}, \text{2-loop}) \simeq -10 \left(\frac{\alpha}{\pi} \right) a_{\mu}(\text{EW}, \text{1-loop}) \left(\ln \frac{m_Z}{m_{\mu}} + 1 \right),$$

where the factor of 10 appears since many "order one" diagrams accidentally add up coherently.

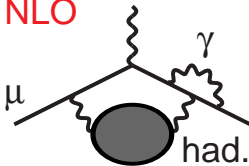
Hadronic Contributions

There are several hadronic contributions:

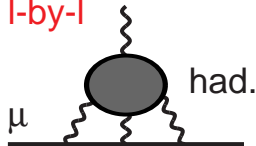
LO



NLO



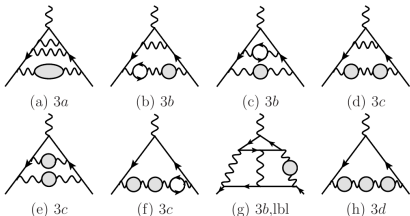
I-by-I



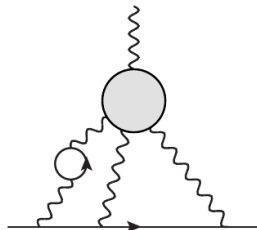
LO: Leading Order (or Vacuum Polarization) Hadronic Contribution

NLO: Next-to-Leading Order Hadronic Contribution

I-by-I: Hadronic light-by-light Contribution



NNLO Hadronic Contributions

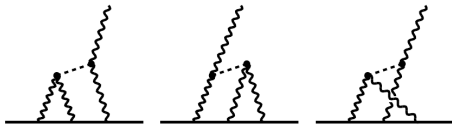


Hadronic I-by-I NLO Contrib.

Modern evaluation of I-by-I contribution

(Melnikov & Vainshtein)

1. First, use the large N_C expansion to find that the leading contribution is the pion pole contribution.

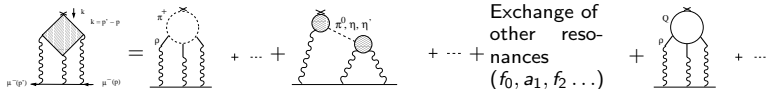


2. Choose the momentum-dependence of the $\pi\gamma\gamma$ coupling (form factor) in such a way that it is consistent with a constraint from QCD (OPE) at the momentum region $q_1^2 \sim q_2^2 \gg q_3^2$. Integrate over the loop momenta.
3. Repeat the above for η, η', a_1, \dots . Basically that's all for the LO in $1/N_C$.
4. As for NLO in $1/N_C$, it depends on authors which diagram is numerically important.

For example,

$$a_\mu^{\text{IbyI}} = \begin{cases} (10.5 \pm 2.6) \times 10^{-10} & \text{'Glasgow consensus', arXiv:0901.0306} \\ (9.8 \pm 2.6) \times 10^{-10} & \text{'G.c.' w/ correction by Nyffeler, PRD94(2016)053006} \\ (10.2 \pm 3.9) \times 10^{-10} & \text{Nyffeler, arXiv:1710.09742} \end{cases}$$

HLbL in muon $g - 2$: summary of selected results (model calculations)



de Rafael '94:

Chiral counting: p^4

N_C -counting: 1

Contribution to $a_{\mu} \times 10^{11}$:

p^6

N_C

p^8

N_C

p^8

N_C

Exchange of other resonances ($f_0, a_1, f_2 \dots$)

BPP:	+83 (32)
HKS:	+90 (15)
KN:	+80 (40)
MV:	+136 (25)
2007:	+110 (40)
PdRV:	+105 (26)
N,JN:	+116 (39)

-19 (13)

-5 (8)

0 (10)

-19 (19)

-19 (13)

ud.: -45

+85 (13)
+83 (6)
+83 (12)
+114 (10)
+114 (13)
+99 (16)

ud.: $+\infty$

-4 (3) [f_0, a_1]

+1.7 (1.7) [a_1]

+22 (5) [a_1]

+8 (12) [f_0, a_1]

+15 (7) [f_0, a_1]

+21 (3)

+10 (11)

0

+2.3 [c-quark]

+21 (3)

ud.: +60

ud. = undressed, i.e. point vertices without form factors

Pseudoscalars: numerically dominant contribution (according to most models !).

Recall (in units of 10^{-11}): $\delta a_{\mu}(\text{HVP}) \approx 40$; $\delta a_{\mu}(\text{exp [BNL]}) = 63$; $\delta a_{\mu}(\text{future exp}) = 16$

BPP = Bijmens, Pallante, Prades '96, '02; HKS = Hayakawa, Kinoshita, Sanda '96, '98, '02; KN = Knecht, AN '02; MV = Melnikov, Vainshtein '04; 2007 = Bijmens, Prades; Miller, de Rafael, Roberts; PdRV = Prades, de Rafael, Vainshtein '09 (compilation; "Glasgow consensus"); N,JN = AN '09; Jegerlehner, AN '09 (compilation)

Recent reevaluations of axial vector contribution lead to much smaller estimates than in MV '04:

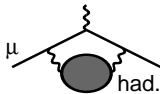
$a_{\mu}^{\text{HLbL}; \text{axial}} = (8 \pm 3) \times 10^{-11}$ (Pauk, Vanderhaeghen '14; Jegerlehner '14, '15). Would shift central values of compilations downwards:

$a_{\mu}^{\text{HLbL}} = (98 \pm 26) \times 10^{-11}$ (PdRV) and $a_{\mu}^{\text{HLbL}} = (102 \pm 39) \times 10^{-11}$ (N, JN).

Slide by A. Nyffeler (Mainz) at 'Muon g-2 lbyl Workshop' at Connecticut, March 12-14, 2018

LO Hadronic Vacuum Polarization Contribution

The diagram to be evaluated:

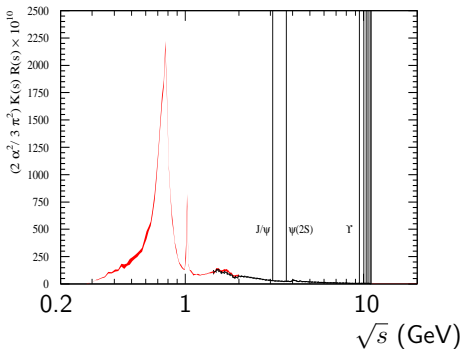


pQCD not useful. Use the **dispersion relation** and the **optical theorem**.

$$\text{had.} = \int \frac{ds}{\pi(s-q^2)} \text{Im had.}$$

$$2 \text{Im had.} = \sum_{\text{had.}} \int d\Phi \left| \text{had.} \right|^2$$

$$a_{\mu}^{\text{had,LO}} = \frac{m_{\mu}^2}{12\pi^3} \int_{s_{\text{th}}}^{\infty} ds \frac{1}{s} \hat{K}(s) \sigma_{\text{had}}(s)$$



- Weight function $\hat{K}(s)/s = \mathcal{O}(1)/s$
 \implies **Lower** energies **more important**
 $\implies \pi^+\pi^-$ channel: 73% of total $a_{\mu}^{\text{had,LO}}$

Main improvements between HLMNT11 and KNT18/19

- Lots of **new input $\sigma(e^+e^- \rightarrow \text{hadrons})$ data**
- Improvements in the estimates of uncertainties due to **radiative corrections** (Vacuum Polarization Radiative Corrections & Final State Radiations)
- Improvements in **data-combination** method

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Channel	Energy range [GeV]	$a_\mu^{\text{had,LOVP}} \times 10^{10}$	$\Delta a_\mu^{(S)}(M_Z^2) \times 10^4$	New data
Chiral perturbation theory (ChPT) threshold contributions				
$\pi^0\gamma$	$m_\pi \leq \sqrt{s} \leq 0.600$	0.12 ± 0.01	0.00 ± 0.00	...
$\pi^+\pi^-$	$2m_\pi \leq \sqrt{s} \leq 0.305$	0.87 ± 0.02	0.01 ± 0.00	...
$\pi^+\pi^-\pi^0$	$3m_\pi \leq \sqrt{s} \leq 0.660$	0.01 ± 0.00	0.00 ± 0.00	...
$\eta\gamma$	$m_\eta \leq \sqrt{s} \leq 0.660$	0.00 ± 0.00	0.00 ± 0.00	...
Data based channels ($\sqrt{s} \leq 1.937$ GeV)				
$\pi^0\gamma$	$0.600 \leq \sqrt{s} \leq 1.350$	4.46 ± 0.10	0.36 ± 0.01	[65]
$\pi^+\pi^-$	$0.305 \leq \sqrt{s} \leq 1.937$	502.97 ± 1.97	34.26 ± 0.12	[34,35]
$\pi^+\pi^-\pi^0$	$0.660 \leq \sqrt{s} \leq 1.937$	47.79 ± 0.89	4.77 ± 0.08	[36]
$\pi^+\pi^-\pi^+\pi^-$	$0.613 \leq \sqrt{s} \leq 1.937$	14.87 ± 0.20	4.02 ± 0.05	[40,42]
$\pi^+\pi^-\pi^0\pi^0$	$0.850 \leq \sqrt{s} \leq 1.937$	19.39 ± 0.78	5.00 ± 0.20	[44]
$(2\pi^+2\pi^-\pi^0)_{\text{non}}$	$1.013 \leq \sqrt{s} \leq 1.937$	0.99 ± 0.09	0.33 ± 0.03	...
$3\pi^+3\pi^-$	$1.313 \leq \sqrt{s} \leq 1.937$	0.23 ± 0.01	0.09 ± 0.01	[66]
$(2\pi^+2\pi^-2\pi^0)_{\text{non}}$	$1.322 \leq \sqrt{s} \leq 1.937$	1.35 ± 0.17	0.51 ± 0.06	...
K^+K^-	$0.988 \leq \sqrt{s} \leq 1.937$	23.03 ± 0.22	3.37 ± 0.03	[45,46,49]
$K_S^0 K_L^0$	$1.004 \leq \sqrt{s} \leq 1.937$	13.04 ± 0.19	1.77 ± 0.03	[50,51]
$KK\pi$	$1.260 \leq \sqrt{s} \leq 1.937$	2.71 ± 0.12	0.89 ± 0.04	[53,54]
$KK2\pi$	$1.350 \leq \sqrt{s} \leq 1.937$	1.93 ± 0.08	0.75 ± 0.03	[50,53,55]
$\eta\gamma$	$0.660 \leq \sqrt{s} \leq 1.760$	0.70 ± 0.02	0.09 ± 0.00	[67]
$\eta\pi^+\pi^-$	$1.091 \leq \sqrt{s} \leq 1.937$	1.29 ± 0.06	0.39 ± 0.02	[68,69]
$(\eta\pi^+\pi^-\pi^0)_{\text{non}}$	$1.333 \leq \sqrt{s} \leq 1.937$	0.60 ± 0.15	0.21 ± 0.05	[70]
$\eta 2\pi^+ 2\pi^-$	$1.338 \leq \sqrt{s} \leq 1.937$	0.08 ± 0.01	0.03 ± 0.00	...
$\eta\omega$	$1.333 \leq \sqrt{s} \leq 1.937$	0.31 ± 0.03	0.10 ± 0.01	[70,71]
$\omega(\rightarrow \pi^0\gamma)\pi^0$	$0.920 \leq \sqrt{s} \leq 1.937$	0.88 ± 0.02	0.19 ± 0.00	[72,73]
$\eta\phi$	$1.569 \leq \sqrt{s} \leq 1.937$	0.42 ± 0.03	0.15 ± 0.01	...
$\phi \rightarrow$ unaccounted	$0.988 \leq \sqrt{s} \leq 1.029$	0.04 ± 0.04	0.01 ± 0.01	...
$\eta\omega\pi^0$	$1.550 \leq \sqrt{s} \leq 1.937$	0.35 ± 0.09	0.14 ± 0.04	[74]
$\eta(\rightarrow \text{npp})K\bar{K}_{\text{non}\phi \rightarrow K\bar{K}}$	$1.569 \leq \sqrt{s} \leq 1.937$	0.01 ± 0.02	0.00 ± 0.01	[53,75]
$p\bar{p}$	$1.890 \leq \sqrt{s} \leq 1.937$	0.03 ± 0.00	0.01 ± 0.00	[76]
$n\bar{n}$	$1.912 \leq \sqrt{s} \leq 1.937$	0.03 ± 0.01	0.01 ± 0.00	[77]
Estimated contributions ($\sqrt{s} \leq 1.937$ GeV)				
$(\pi^+\pi^-3\pi^0)_{\text{non}}$	$1.013 \leq \sqrt{s} \leq 1.937$	0.50 ± 0.04	0.16 ± 0.01	...
$(\pi^+\pi^-4\pi^0)_{\text{non}}$	$1.313 \leq \sqrt{s} \leq 1.937$	0.21 ± 0.21	0.08 ± 0.08	...
$KK3\pi$	$1.569 \leq \sqrt{s} \leq 1.937$	0.03 ± 0.02	0.02 ± 0.01	...
$\omega(\rightarrow \text{npp})2\pi$	$1.285 \leq \sqrt{s} \leq 1.937$	0.10 ± 0.02	0.03 ± 0.01	...
$\omega(\rightarrow \text{npp})3\pi$	$1.322 \leq \sqrt{s} \leq 1.937$	0.17 ± 0.03	0.06 ± 0.01	...
$\omega(\rightarrow \text{npp})KK$	$1.569 \leq \sqrt{s} \leq 1.937$	0.00 ± 0.00	0.00 ± 0.00	...
$\eta\pi^+\pi^-2\pi^0$	$1.338 \leq \sqrt{s} \leq 1.937$	0.08 ± 0.04	0.03 ± 0.02	...
Other contributions ($\sqrt{s} > 1.937$ GeV)				
Inclusive channel	$1.937 \leq \sqrt{s} \leq 11.199$	43.67 ± 0.67	82.82 ± 1.05	[56,62,63]
J/ψ	...	6.26 ± 0.19	7.07 ± 0.22	...
ψ'	...	1.58 ± 0.04	2.51 ± 0.06	...
$\Upsilon(1S-4S)$...	0.09 ± 0.00	1.06 ± 0.02	...
pQCD	$11.199 \leq \sqrt{s} \leq \infty$	2.07 ± 0.00	124.79 ± 0.10	...
Total	$m_\pi \leq \sqrt{s} \leq \infty$	693.26 ± 2.46	276.11 ± 1.11	...

Breakdown of contributions to a_μ (had, LO VP) from various hadronic final states

We have included new data sets from ~ 30 papers, in addition to those included in the HLMNT11 analysis

We have included ~ 30 hadronic final states

At $2 \lesssim \sqrt{s} \lesssim 11$ GeV, we use inclusively measured data

At higher energies $\gtrsim 11$ GeV, we use pQCD

Main improvements between HLMNT11 and KNT18/19

- Lots of new input $\sigma(e^+e^- \rightarrow \text{hadrons})$ data
- Improvements in the estimates of uncertainties due to **radiative corrections** (Vacuum Polarization Radiative Corrections & Final State Radiations)
- Improvements in data-combination method

Vacuum Polarization Corrections to $\sigma(e^+e^- \rightarrow \text{hadrons})$

Optical Theorem:

$$\text{Im} \left| \begin{array}{c} \gamma \\ \text{had} \\ \gamma \end{array} \right| \Leftrightarrow \left| \begin{array}{c} \gamma \\ \text{had} \\ \gamma \end{array} \right|^2$$

$\text{Im} \Pi_{\text{had}}(q^2) \qquad \sim \sigma_{\text{had}}(q^2)$

Experimentally observed cross section:

$$\begin{array}{c} e^+ \\ \swarrow \\ \gamma \\ \swarrow \\ e^- \end{array} \begin{array}{c} \text{VP} \\ \text{had} \end{array} \Rightarrow \begin{array}{c} e^+ \\ \swarrow \\ \gamma \\ \swarrow \\ e^- \end{array} \begin{array}{c} \text{had} \end{array}$$

$\sigma(e^+e^- \rightarrow \gamma^* \rightarrow \text{hadrons}) \qquad \sigma^0(e^+e^- \rightarrow \gamma^* \rightarrow \text{hadrons})$

To evaluate $a_{\mu}^{\text{LO, had}}$, we need to subtract the vacuum polarization (VP) contribution.

It is straightforward to subtract the leptonic part of the VP, but the **hadronic part is non-trivial**: we need to do this **recursively** by using hadronic data. (We did this in the KNT18 paper.)

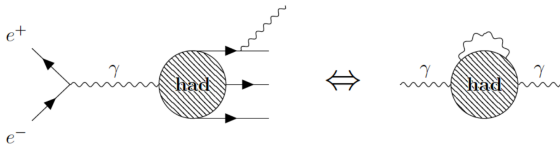
Final State Radiation Corrections to $\sigma(e^+e^- \rightarrow \text{hadrons})$

Optical Theorem:

$$\text{Im} \left| \begin{array}{c} \gamma \\ \text{had} \\ \gamma \end{array} \right| \Leftrightarrow \left| \begin{array}{c} \gamma \\ \text{had} \\ \gamma \end{array} \right|^2$$

$\text{Im} \Pi_{\text{had}}(q^2) \qquad \qquad \qquad \sim \sigma_{\text{had}}(q^2)$

To evaluate $a_{\mu}^{\text{LO, had}}$, by definition, we use the hadronic cross sections which include all the Final State Radiations (FSR).



In real experiments, people often impose cuts on the final state photons and/or miss photons in the final states. So we have to add back those missed photons, which introduces uncertainties.

In KNT18, we revisited the FSR corrections in the K^+K^- and $K_S^0K_L^0$ final states, and found smaller FSR uncertainties than our previous papers.

Main improvements between HLMNT11 and KNT18/19

- Lots of new input $\sigma(e^+e^- \rightarrow \text{hadrons})$ data
- Improvements in the estimates of uncertainties due to radiative corrections (Vacuum Polarization Radiative Corrections & Final State Radiations)
- Improvements in **data-combination** method

Data Combination

To evaluate the vacuum polarization contribution, we have to combine lots of experimental data.

To do so, we usually construct a χ^2 function and find the value of $R(s)$ at each bin which minimizes χ^2 .

Naively, the χ^2 function defined as

$$\chi^2(\{\bar{R}_i\}) \equiv \sum_{n=1}^{N_{\text{exp}}} \sum_{i=1}^{N_{\text{bin}}} \sum_{j=1}^{N_{\text{bin}}} (R_i^{(n)} - \bar{R}_i)(V_n^{-1})_{ij}(R_j^{(n)} - \bar{R}_j),$$

where V_n is the cov. matrix of the n -th exp.,

$$V_{n,ij} = \begin{cases} (\delta R_{i,\text{stat}}^{(n)})^2 + (\delta R_{i,\text{sys}}^{(n)})^2 & (\text{for } i = j) \\ (\delta R_{i,\text{sys}}^{(n)})(\delta R_{j,\text{sys}}^{(n)}) & (\text{for } i \neq j) \end{cases}$$

may seem OK, but when there are non-negligible normalization uncertainties in the data, we have to be more careful.

χ^2 vs normalization error: d'Agostini bias

G. D'Agostini, Nucl. Instrum. Meth. A346 (1994) 306

We first consider an observable x whose true value is 1.
Suppose that there is an experiment which measures x
and whose normalization uncertainty is 10%.

Now, assume that this experiment measured x twice:

$$\text{1st result: } 0.9 \pm 0.1_{\text{stat}} \pm 10\%_{\text{syst}} ,$$

$$\text{2nd result: } 1.1 \pm 0.1_{\text{stat}} \pm 10\%_{\text{syst}} .$$

Taking the systematic errors 0.09 and 0.11, respectively,
the covariance matrix and the χ^2 function are

$$(\text{cov.}) = \begin{pmatrix} 0.1^2 + 0.09^2 & 0.09 \cdot 0.11 \\ 0.09 \cdot 0.11 & 0.1^2 + 0.11^2 \end{pmatrix} ,$$

$$\chi^2 = (x - 0.9 \quad x - 1.1) (\text{cov.})^{-1} \begin{pmatrix} x - 0.9 \\ x - 1.1 \end{pmatrix} .$$

χ^2 takes its minimum at $x = 0.98$: **Biased downwards!**

d'Agostini bias (2): improvement by iterations

What was wrong? In the previous page,

$$\text{1st result: } 0.9 \pm 0.1_{\text{stat}} \pm 10\%_{\text{syst}} ,$$

$$\text{2nd result: } 1.1 \pm 0.1_{\text{stat}} \pm 10\%_{\text{syst}} .$$

we took the syst. errors 0.09 and 0.11, respectively, which made the downward bias. Instead, we should take 10% of some estimator \bar{x} as the syst. errors. Then,

$$(\text{cov.}) = \begin{pmatrix} 0.1^2 + (0.1\bar{x})^2 & (0.1\bar{x})^2 \\ (0.1\bar{x})^2 & 0.1^2 + (0.1\bar{x})^2 \end{pmatrix} ,$$

$$\chi^2 = (x - 0.9 \quad x - 1.1) (\text{cov.})^{-1} \begin{pmatrix} x - 0.9 \\ x - 1.1 \end{pmatrix} .$$

χ^2 takes its minimum at $x = 1.00$: **Unbiased!**

In more general cases, we use **iterations**: we find an estimator for the next round of iteration by

χ^2 -minimization.

R.D.Ball et al, JHEP 1005 (2010) 075.

$\sigma(e^+e^- \rightarrow \pi^+\pi^-)$ data

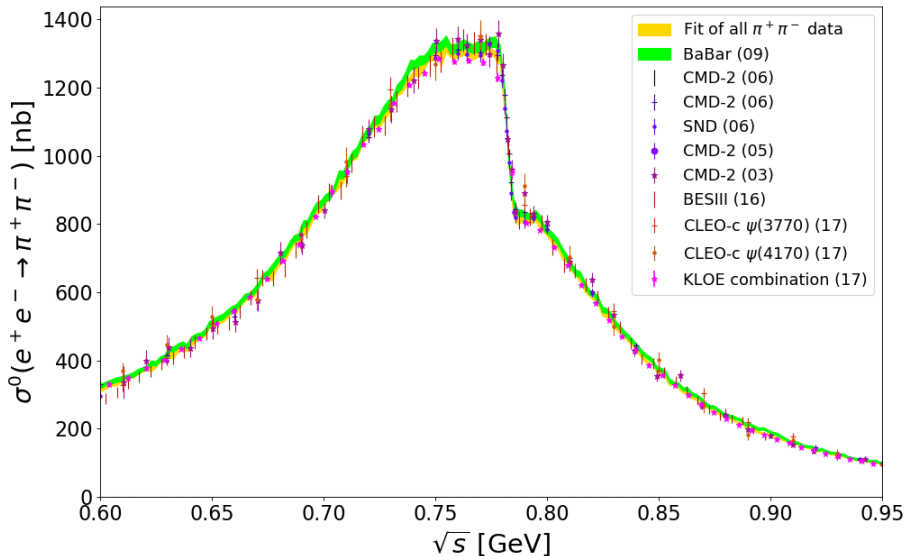


Fig. from KNT19, arXiv:1911.00367

$\sigma(e^+e^- \rightarrow \pi^+\pi^-)$: ρ - ω interference region

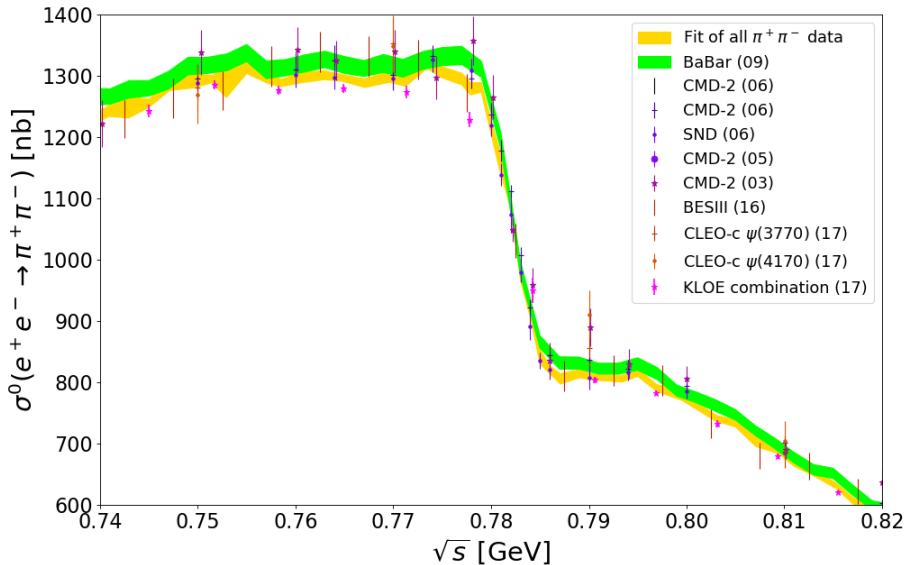


Fig. from KNT19, arXiv:1911.00367

$\sigma(e^+e^- \rightarrow \pi^+\pi^-)$: relative differences

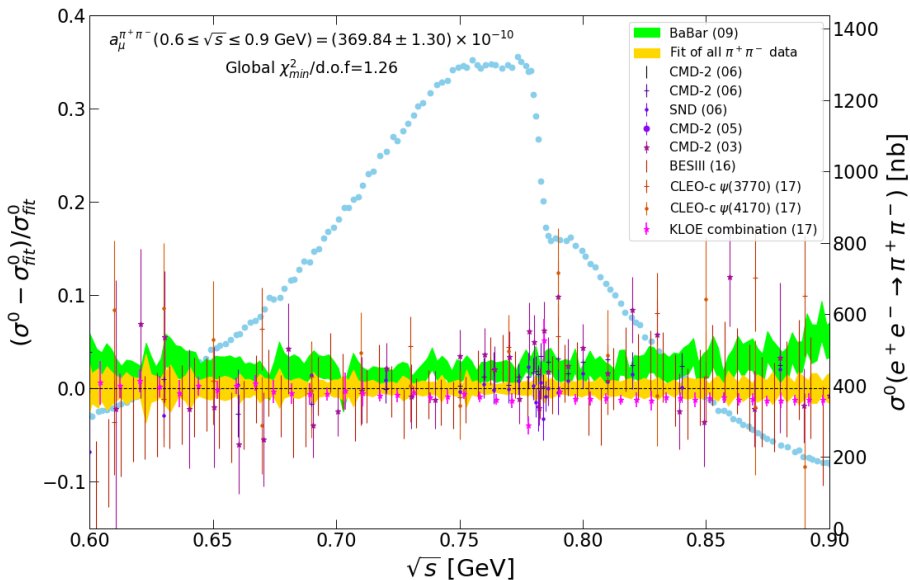


Fig. from KNT19, arXiv:1911.00367

Contribution to $(g - 2)_\mu$ from $\pi^+\pi^-$ channel

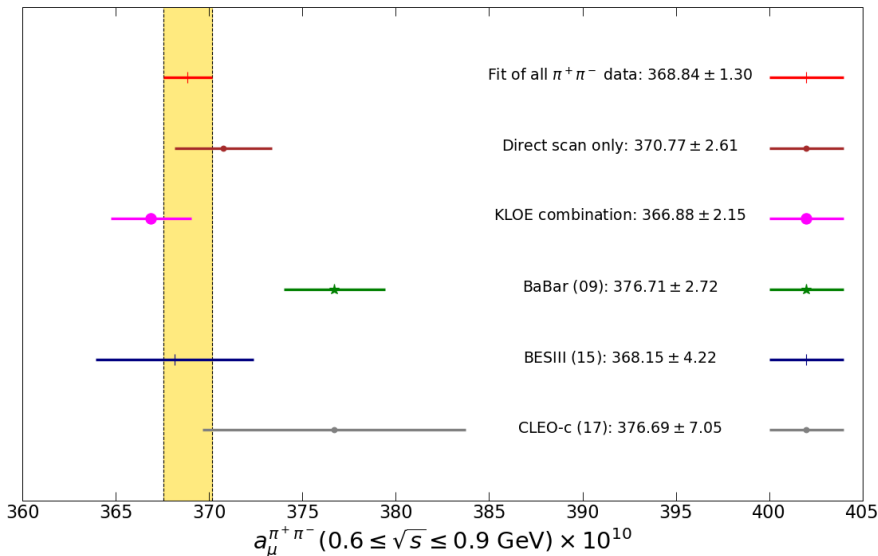
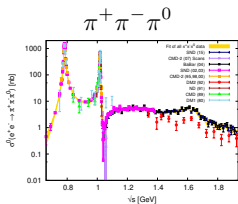
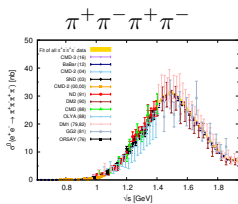


Fig. from KNT19, arXiv:1911.00367

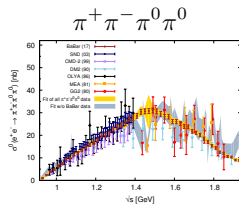
Other notable exclusive channels [KNT18: arXiv:1802.02995, PRD (in press)]



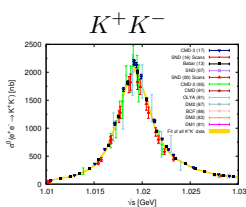
HLMNT11: 47.51 ± 0.99
 KNT18: 47.92 ± 0.89



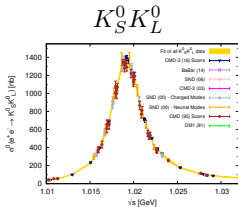
HLMNT11: 14.65 ± 0.47
 KNT18: 14.87 ± 0.20



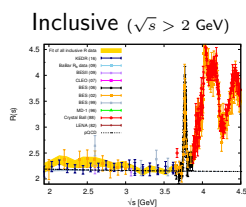
HLMNT11: 20.37 ± 1.26
 KNT18: 19.39 ± 0.78



HLMNT11: 22.15 ± 0.46
 KNT18: 23.03 ± 0.22



HLMNT11: 13.33 ± 0.16
 KNT18: 13.04 ± 0.19



HLMNT11: 41.40 ± 0.87
 KNT18: 41.27 ± 0.62

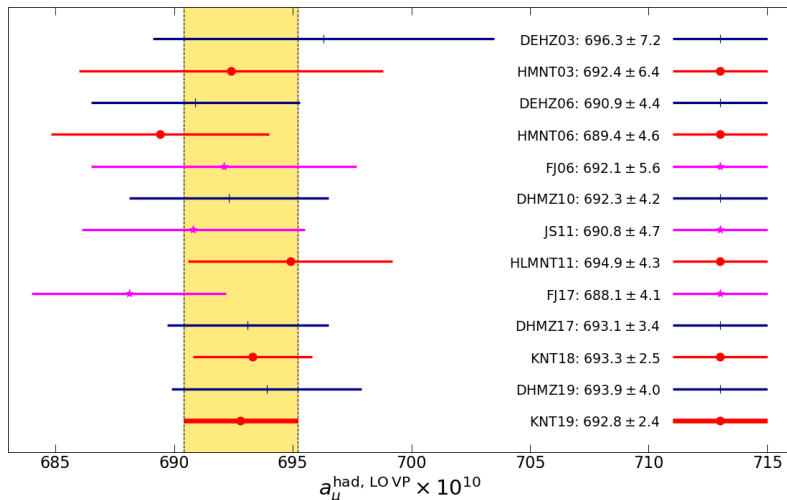
Slide by A. Keshavarzi (Liverpool) at 'Muon $g - 2$ Workshop' at Mainz, June 18-22, 2018

Hadronic VP Contributions: comparison

Adding up all the channels, pQCD & narrow resonances contributions, we get

$$a_{\mu}^{\text{had, LO VP}}(\text{KNT19}) = (692.8 \pm 2.4) \times 10^{-10} \quad (\text{KNT18: } (693.3 \pm 2.5) \times 10^{-10})$$

$$a_{\mu}^{\text{had, NLO VP}}(\text{KNT19}) = (-9.83 \pm 0.04) \times 10^{-10} \quad (\text{KNT18: } (-9.82 \pm 0.04) \times 10^{-10})$$



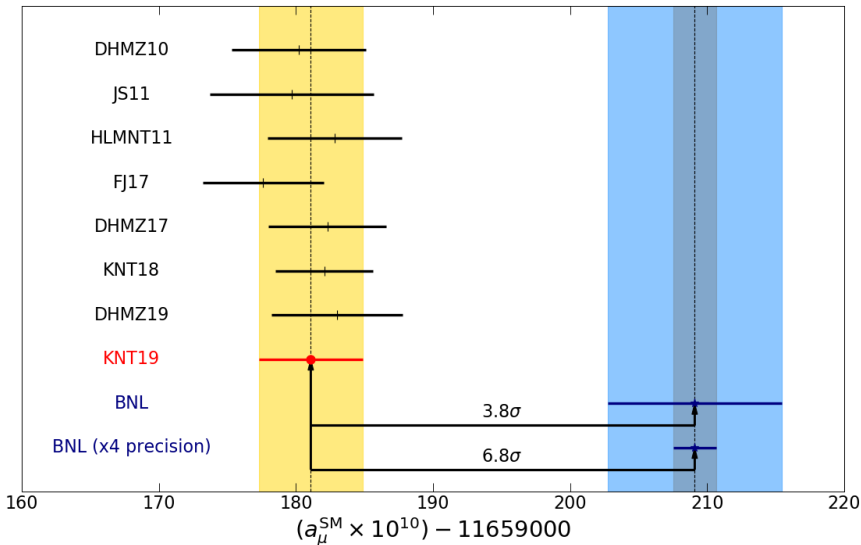
Breakdown of SM prediction for muon g-2

	<u>2011</u>	→	<u>2018</u>	<u>2019</u>
QED	11658471.81 (0.02)	→	11658471.90 (0.01) [arXiv:1712.06060]	
EW	15.40 (0.20)	→	15.36 (0.10) [Phys. Rev. D 88 (2013) 053005]	
LO HLbL	10.50 (2.60)	→	9.80 (2.60) [EPJ Web Conf. 118 (2016) 01005]	9.34 (2.92)
NLO HLbL			0.30 (0.20) [Phys. Lett. B 735 (2014) 90]	
	<u>HLMNT11</u>		<u>KNT18</u>	<u>KNT19</u>
LO HVP	694.91 (4.27)	→	693.27 (2.46) this work	692.78 (2.42)
NLO HVP	-9.84 (0.07)	→	-9.82 (0.04) this work	-9.83 (0.04)
NNLO HVP			1.24 (0.01) [Phys. Lett. B 734 (2014) 144]	
Theory total	11659182.80 (4.94)	→	11659182.05 (3.56) this work...	181.08 (3.78)
Experiment			11659209.10 (6.33) world avg	
Exp - Theory	26.1 (8.0)	→	27.1 (7.3) this work	28.0 (7.4)
Δa_μ	3.3 σ	→	3.7 σ this work	3.8 σ

(HVP: Hadronic Vacuum Polarization)
(HLbL: Hadronic Light-by-Light)

(Numbers taken from KNT18
and from KNT19)

Exp. value of muon g-2 vs SM prediction



Comparison with Other Work

Contributions from major channels to $a_\mu(\text{LO,had})$ for $\sqrt{s} < 1.8\text{GeV}$:

channel	KNT18	DHMZ19	diff
$\pi^+\pi^-$	503.74 ± 1.96	507.80 ± 3.35	-4.06
$\pi^+\pi^-\pi^0$	47.70 ± 0.89	46.20 ± 1.45	1.50
K^+K^-	23.00 ± 0.22	23.08 ± 0.44	-0.08
$\pi^+\pi^-2\pi^0$	18.15 ± 0.74	18.01 ± 0.55	0.14
$2\pi^+2\pi^-$	13.99 ± 0.19	13.68 ± 0.31	0.31
$K_S^0K_L^0$	13.04 ± 0.19	12.82 ± 0.24	0.22
$\pi^0\gamma$	4.58 ± 0.10	4.29 ± 0.10	0.29
\vdots	\vdots	\vdots	\vdots

“DHMZ19” = M. Davier et al, arXiv:1908.00921

Difference in the $\pi^+\pi^-$ channel is mainly from the way to combine the data sets.

KNT18: Global χ^2 minimization

DHMZ19: Takes the average of “all but KLOE” and “all but BaBar” as the mean value, and counts the half of the diff of the two as an additional systematic uncertainty.

Comparison with Lattice Results

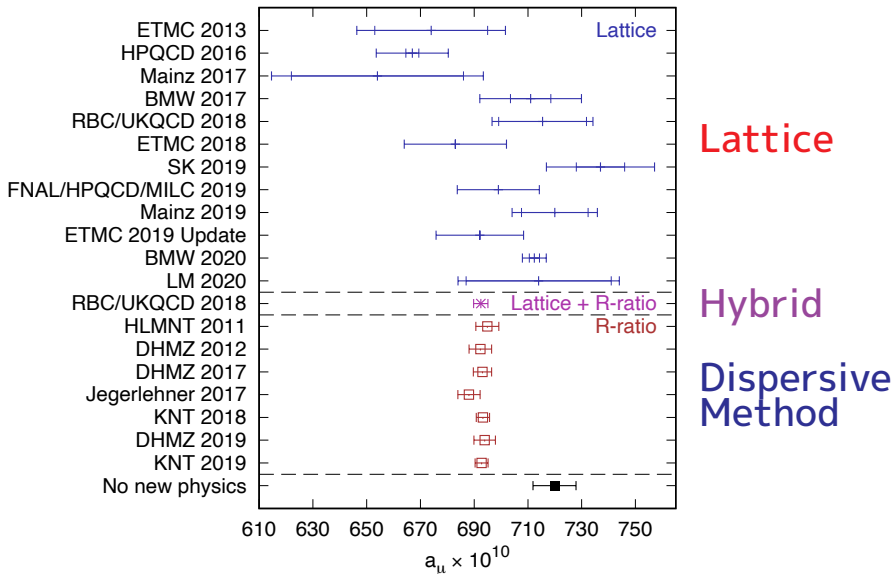


Fig. from C. Lehner & A. Meyer, arXiv:2003.04177

Muon g-2 Theory Initiative

Steering Committee:

G. Colangelo	(Hadron Theory)	HVP and HLbL
M. Davier	(e^+e^- exp. (BaBar))	HVP
S. Eidelman	(e^+e^- exp. (CMD-2, CMD-3 & SND))	HVP
A. El-Khadra	(Lattice QCD)	HVP
C. Lehner	(Lattice QCD)	HVP and HLbL
T. Mibe	(J-PARC g-2 exp.)	
A. Nyffeler	(Hadron Theory)	HLbL
B. L. Roberts	(Fermilab g-2 exp.)	
T. Teubner	(Hadron Theory)	HVP

HVP: Hadronic Vacuum Polarization

HLbL: Hadronic Light-by-Light

Muon g-2 Theory Initiative: Goals

- ① theory support to the Fermilab and J-PARC experiments to maximize their impact
 - ⇒ need theoretical predictions of the hadronic corrections with reduced and reliably estimated uncertainties
- ② summarize the theoretical calculations of the hadronic corrections to the muon g-2
 - ⇒ comparisons of intermediate quantities between the different approaches. For example, lattice vs experiment
 - ⇒ assess reliability of uncertainty estimates
- ③ combine to provide theory predictions for a_{μ}^{HVP} and a_{μ}^{HLbL} and write a report **before** the Fermilab and J-PARC experiments announce their first results.

slide by A. El-Khadra at Phipsi17, June 26-29, 2017

(Underlines by DN)

Muon g-2 Theory Initiative: Workshops

- 1st plenary workshop: near Fermilab, June 2017
- Hadronic Vacuum Polarization workshop: KEK, February 2018
- Hadronic Light-by-Light workshop: Connecticut, March 2018
- 2nd plenary workshop: Mainz, June 2018
- 3rd plenary workshop: Seattle, September 2019
- 4th plenary workshop: KEK, June 2020
→ postponed for fall 2020

We have discussed a lot about the **White Paper**:
In particular, how to come up with **a single theory prediction** to be compared with the exp. result.

Timeline for the White Paper

- 🕒 Earliest possible release date for Fermilab g-2 measurement:
15-20 December 2019
- 🕒 Post the WP on arXiv by: **1 Dec. 2019** → **Submission to arXiv:
Very soon (probably by the end of May)**
- 🕒 Deadline for finalizing individual WP chapters:
1 Nov 2019
At this date the Overleaf chapters will be frozen.
- 🕒 Editorial board will release complete WP to authors for feedback on:
15 Nov. 2019
will need to receive feedback from authors within a week
- 🕒 Experimental and theoretical inputs used in WP must be published by:
15 Oct 2019
To make sure to be included in WP discussion, a paper to be posted in arXiv by same date.

Note: The WP will be posted on arXiv in December, even if the Fermilab experiment's release date is delayed.

slide by A. El-Khadra at the Seattle muon g-2 workshop, September 9-13, 2019

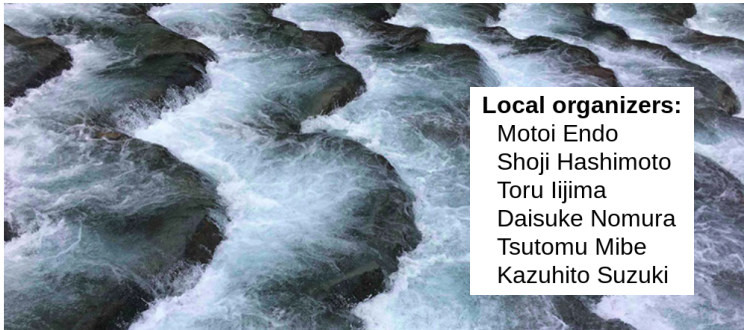
White Paper Outline

- 🌟 Executive Summary
- 🌟 Introduction
- 🌟 Chapter 1: data-driven HVP
- 🌟 Chapter 2: lattice HVP
- 🌟 Chapter 3: data-driven HLbL
- 🌟 Chapter 4: lattice HLbL
- 🌟 Chapter 5: QED + EW
T. Aoyama, T. Kinoshita, M. Nio
D. Stöckinger, H. Stöckinger-Kim
- 🌟 Summary, Conclusions, and Outlook

slide by A. El-Khadra at the Seattle muon g-2 workshop, September 9-13, 2019

Muon g-2 theory initiative workshop

fall
~~June 1-5, 2020~~
KEK, Tsukuba, Japan



Local organizers:

Motoi Endo
Shoji Hashimoto
Toru Iijima
Daisuke Nomura
Tutomu Mibe
Kazuhiro Suzuki

Home

g-2 theory initiative

Registration

Access

Accommodation

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About

The muon g-2 is arguably one of the most important observables in contemporary particle physics. The long-standing anomaly at the level of more than 3 standard deviations between the experimental value

~~June 1-5, 2020~~ at KEK postponed for fall 2020
an activity of the Muon g-2 Theory Initiative

Summary

- **Standard Model prediction for $(g - 2)_\mu$: $\gtrsim 3.5\sigma$ deviation** from measured value \implies New Physics?
- Recent data-driven evaluations of hadronic vacuum polarization contributions seem convergent
- To better establish the $g - 2$ anomaly, better data for $e^+e^- \rightarrow \pi^+\pi^-$ welcome (from CMD-3, SND, Belle II, ...)
- Lattice calculations still suffer from large uncertainties (but a hybrid approach is useful)
- New exp. at Fermilab and J-PARC expected to reduce the uncertainty of $(g - 2)_\mu$ by a factor of 4